

CNGS - CERN Neutrinos to Gran Sasso



> "Neutrinos" and "Gran Sasso"

> Main components, layout at CERN,
"magnetic horns",
expected CNGS beam performance

> Status of works - Schedule

> Summary



A sincere " THANK YOU ! "

to the many colleagues who are contributing,
at CERN and elsewhere, to the CNGS project

- special thanks: Francesco Pietropaolo (INFN / CERN)
- Jean-Luc Caron (AC-DI-MultiMedia)

What are Neutrinos (ν) ?

⇒ elementary particles

⇒ come in three flavors (LEP)
(pistachio, chocolate, vanilla)



⇒ electric charge: zero !

⇒ **mass**: very small **zero?**

⇒ interaction with matter: "very weak"

" the elusive particle "

Leptons

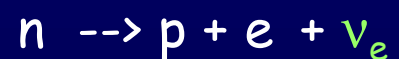
particle	electric charge
e	-1
ν_e	0
<hr/>	
μ	-1
ν_μ	0
<hr/>	
τ	-1
ν_τ	0

+ antiparticles

Where are the Neutrinos ?

⇒ “all around us”

-> radioactive decay of atomic nuclei (e.g. in granite)



-> nuclear reactors

-> from the sun

-> at accelerators... (high energy neutrinos)

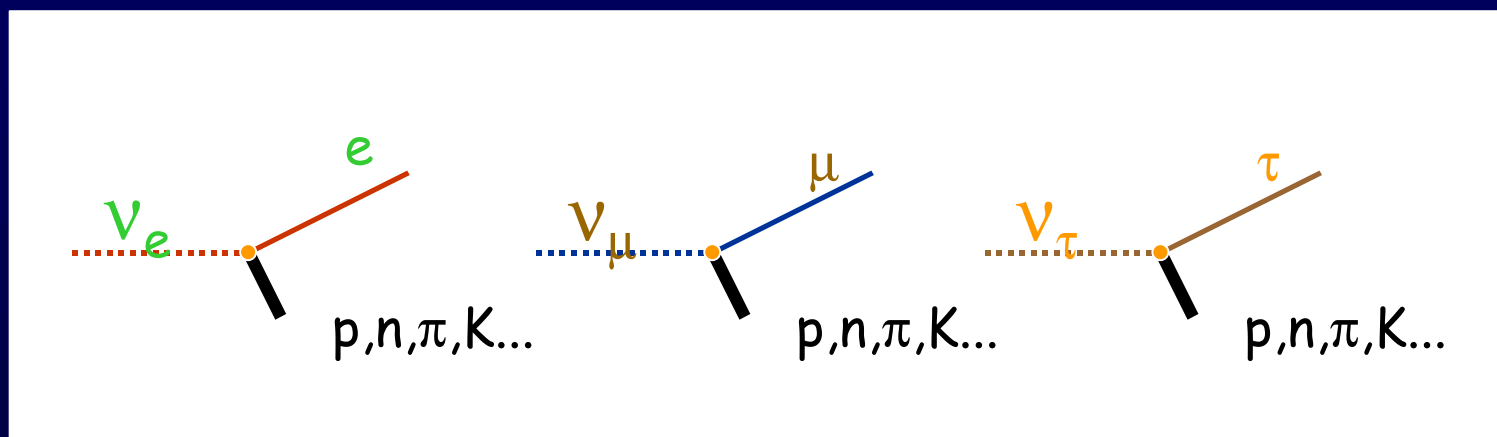


-> from reactions of cosmic rays in the atmosphere

....

How do we detect neutrinos ?

neutrinos interact **VERY** rarely with matter - when they do, they often produce a lepton of their "own character":



NOTE: a minimum amount of energy is needed
(to create the mass of the lepton):

$$m_e = 0.5 \text{ MeV}, \quad - \quad m_\mu = 106 \text{ MeV} \quad - \quad m_\tau = 1770 \text{ MeV}$$

The higher the neutrino energy, the more likely the interaction !

Neutrino mass ?

⇒ Standard model of particle physics: ν masses "ZERO"

⇒ "Direct" mass measurements → upper limits

(in decay experiments measuring kinetic energy of "the partner")

$$m_{\nu_e} < 5 \text{ eV} \quad m_{\nu_\mu} < 170 \text{ keV} \quad m_{\nu_\tau} < 18 \text{ MeV}$$

⇒ What's the problem ?

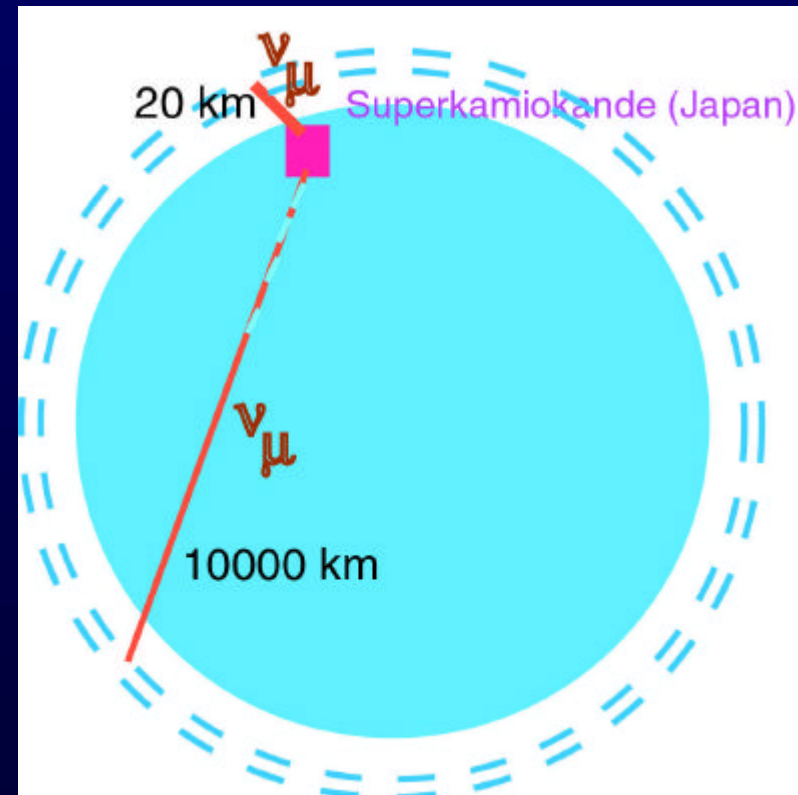
⇒ OBSERVATION 1 : SOLAR NEUTRINO "DEFICIT"

only about 50% of the ν_e expected are actually observed: ν_e disappear
"en route" from the sun to the earth ...

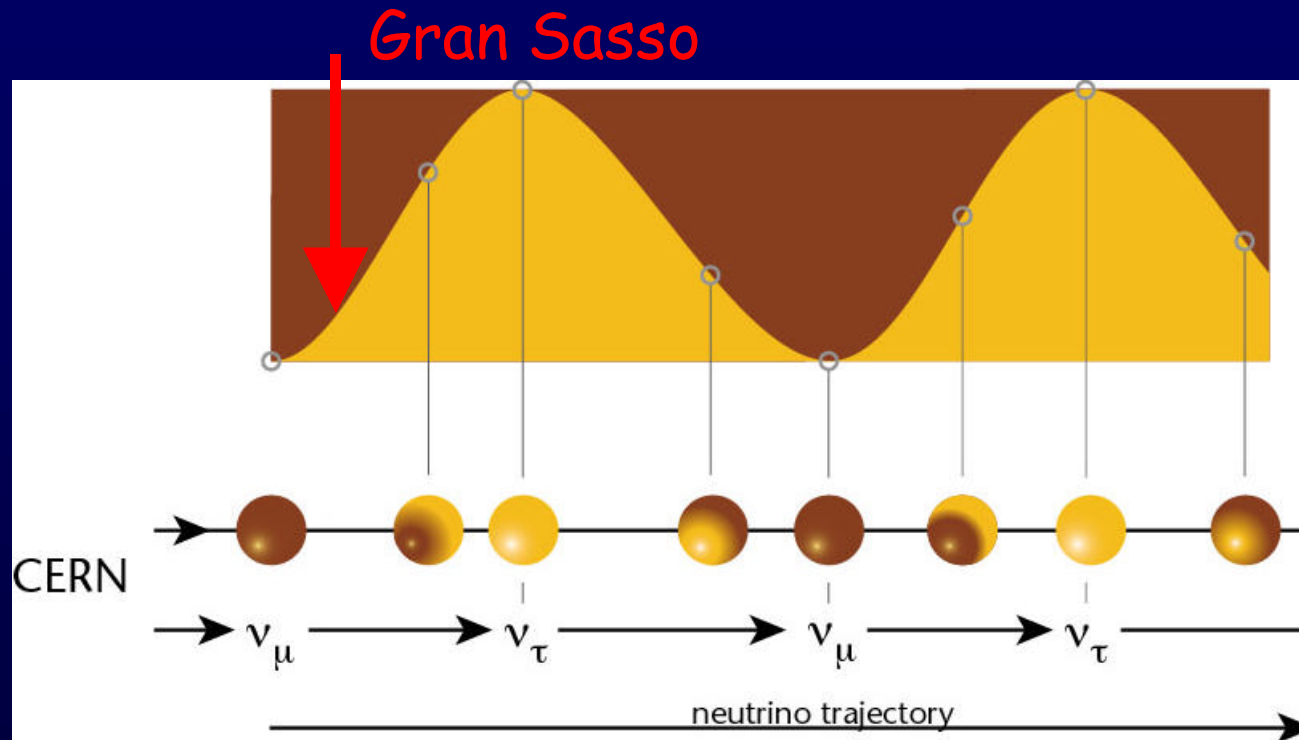
OBSERVATION 2 : "ATMOSPHERIC NEUTRINO ANOMALY"

much less ν_μ "from below"
observed w.r.t. expectations

ν_μ disappear "en route"
over 10'000 km ... ?



Neutrino Oscillation



ν 's change flavor ! Is this possible?

--> Yes, "if neutrinos have mass"!

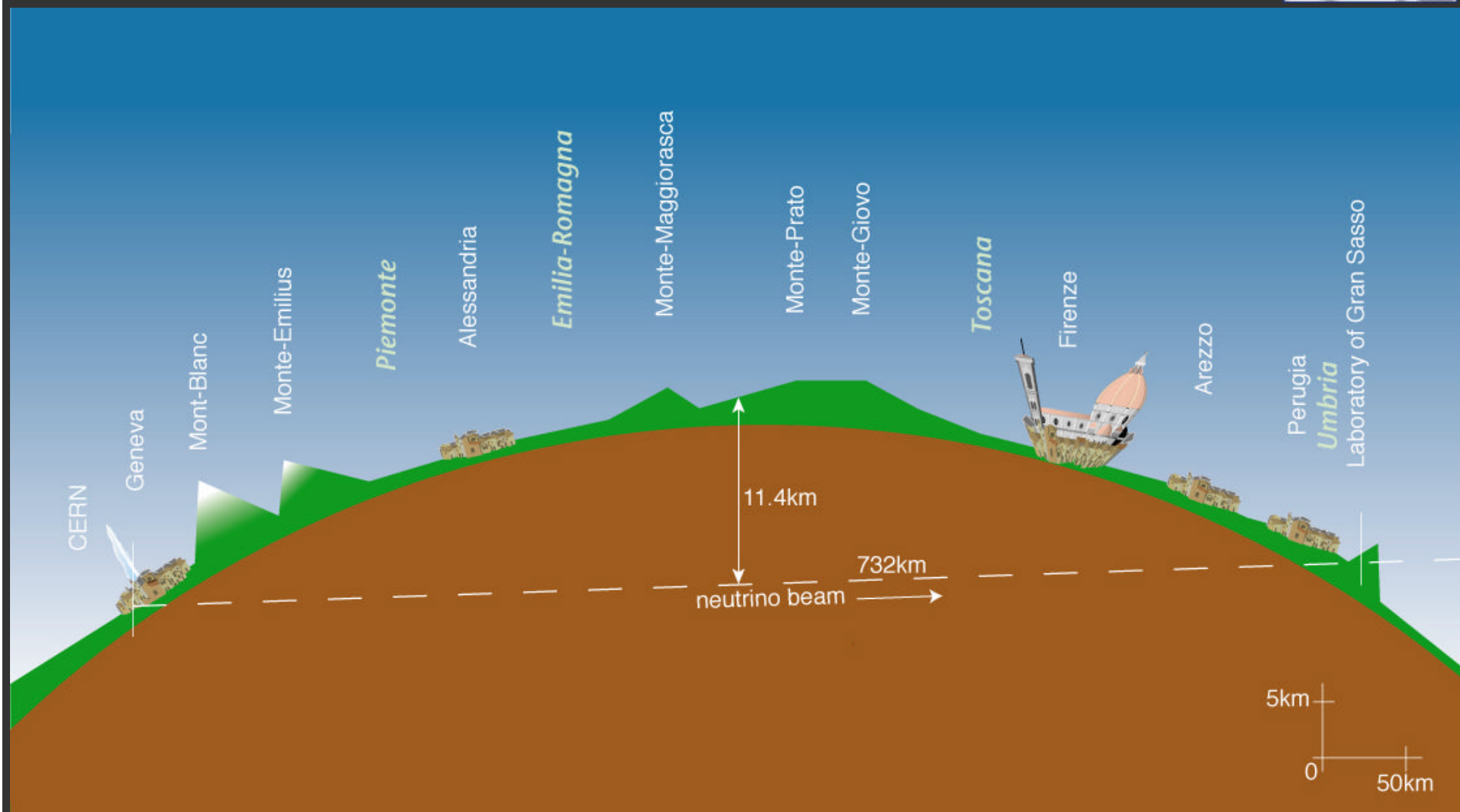


In Dec. 1999, CERN council approved the CNGS project:

- build an intense ν_μ beam at CERN-SPS
- search for ν_τ appearance at Gran Sasso laboratory
(730 km from CERN)

“long base-line” ν_μ -- ν_τ oscillation experiment

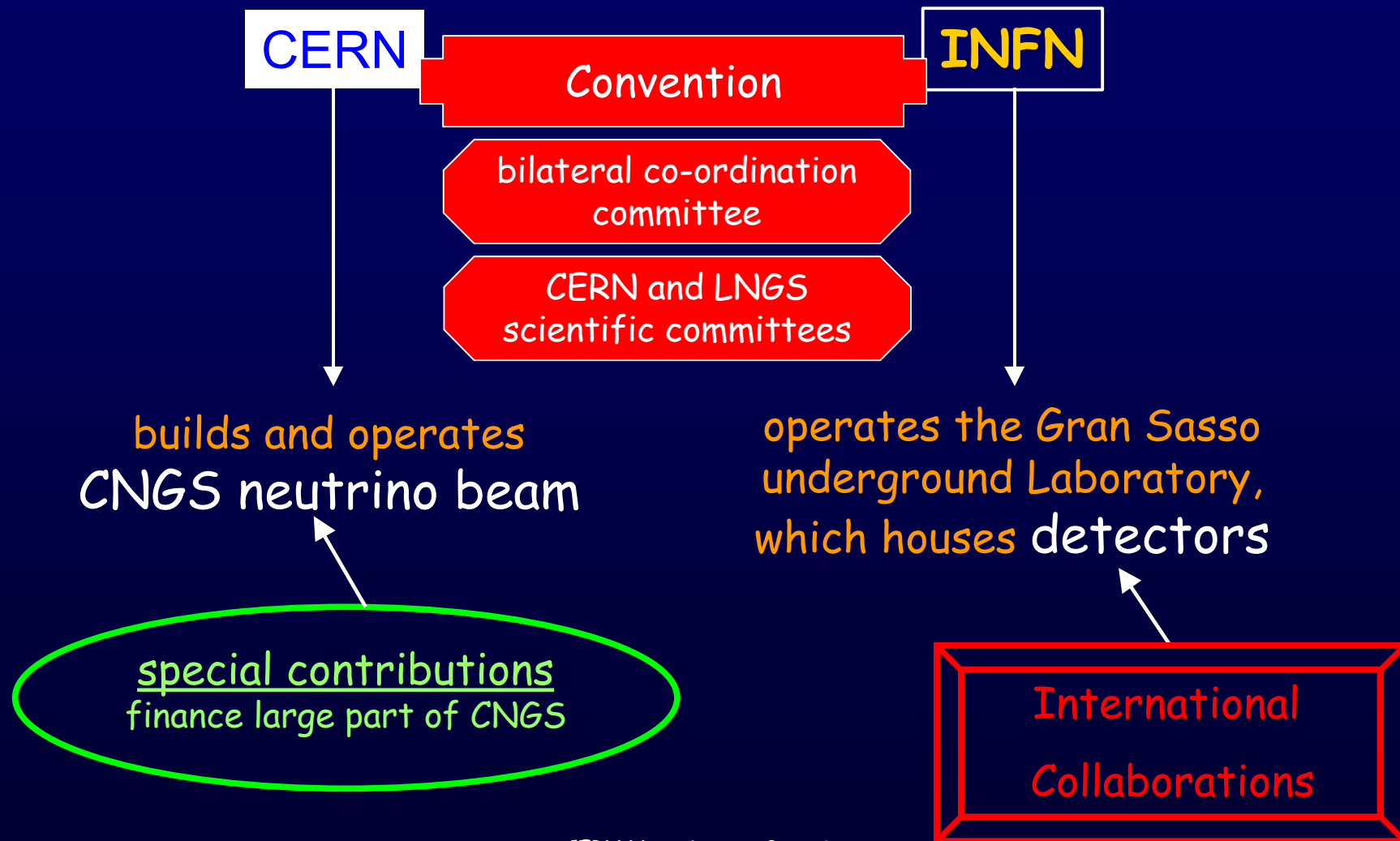
note: K2K (Japan) running; NuMI/MINOS (US) under construction



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ORGANISATION of CNGS

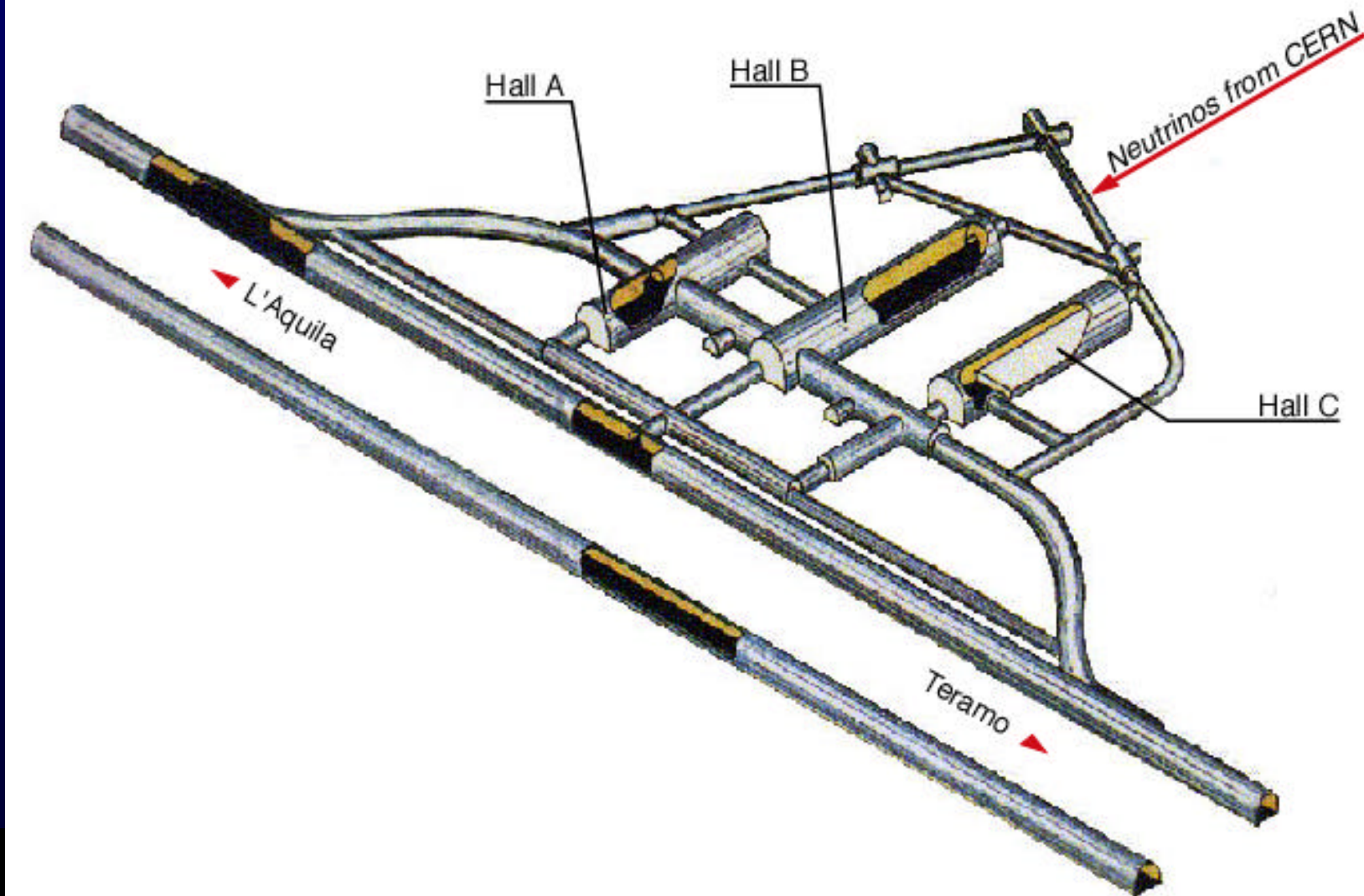




The Gran Sasso Laboratory (LNGS)



The Gran Sasso Laboratory (LNGS)

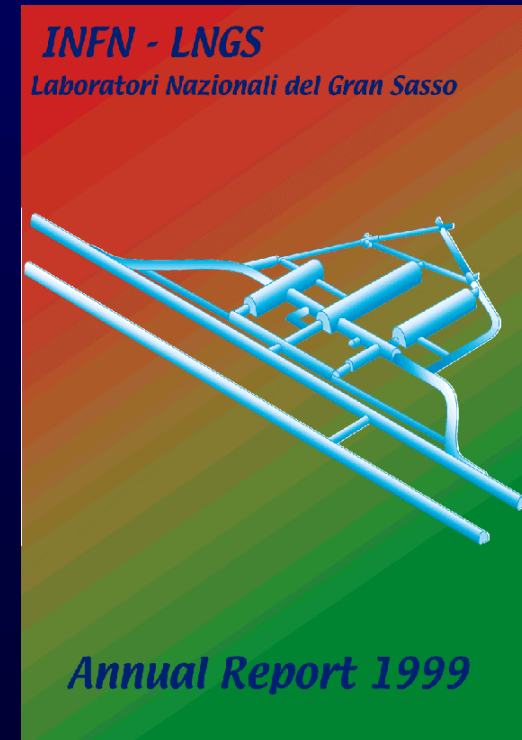
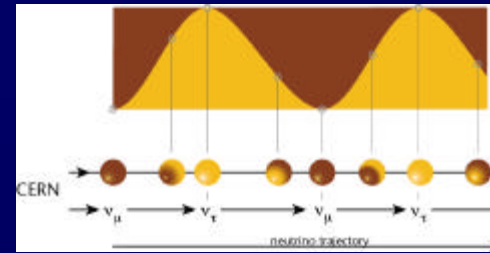


730 km might seem too short -
but look at the details :

Background low enough,
event rate still acceptable
--> 730 km almost perfect

AND, VERY IMPORTANT:

- existing laboratory with
its infrastructure (since 1987)
- large halls directed to CERN
- caverns in the GS mountains:
1500 m of rock shielding
- tradition in very successful
neutrino physics experiments
(solar ν 's)



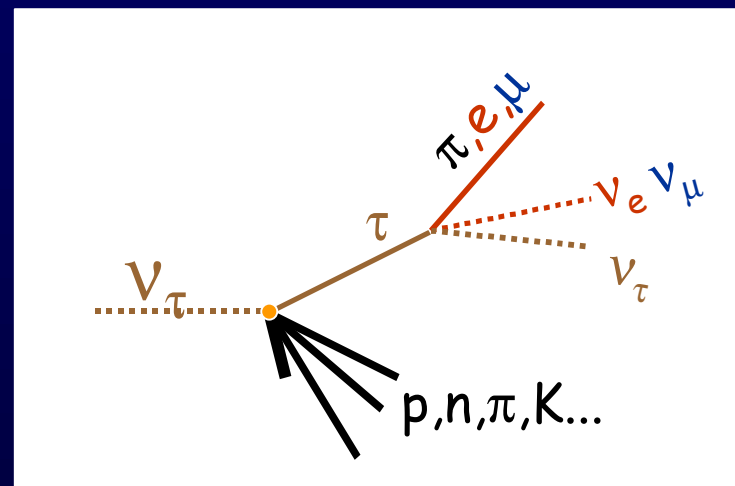
Detecting ν_τ at Gran Sasso



-> look for the τ lepton:

extremely difficult -

τ travels only about 1 mm before decaying



-> two approaches:

(a) very good position resolution (see the decay "kink") -> OPERA

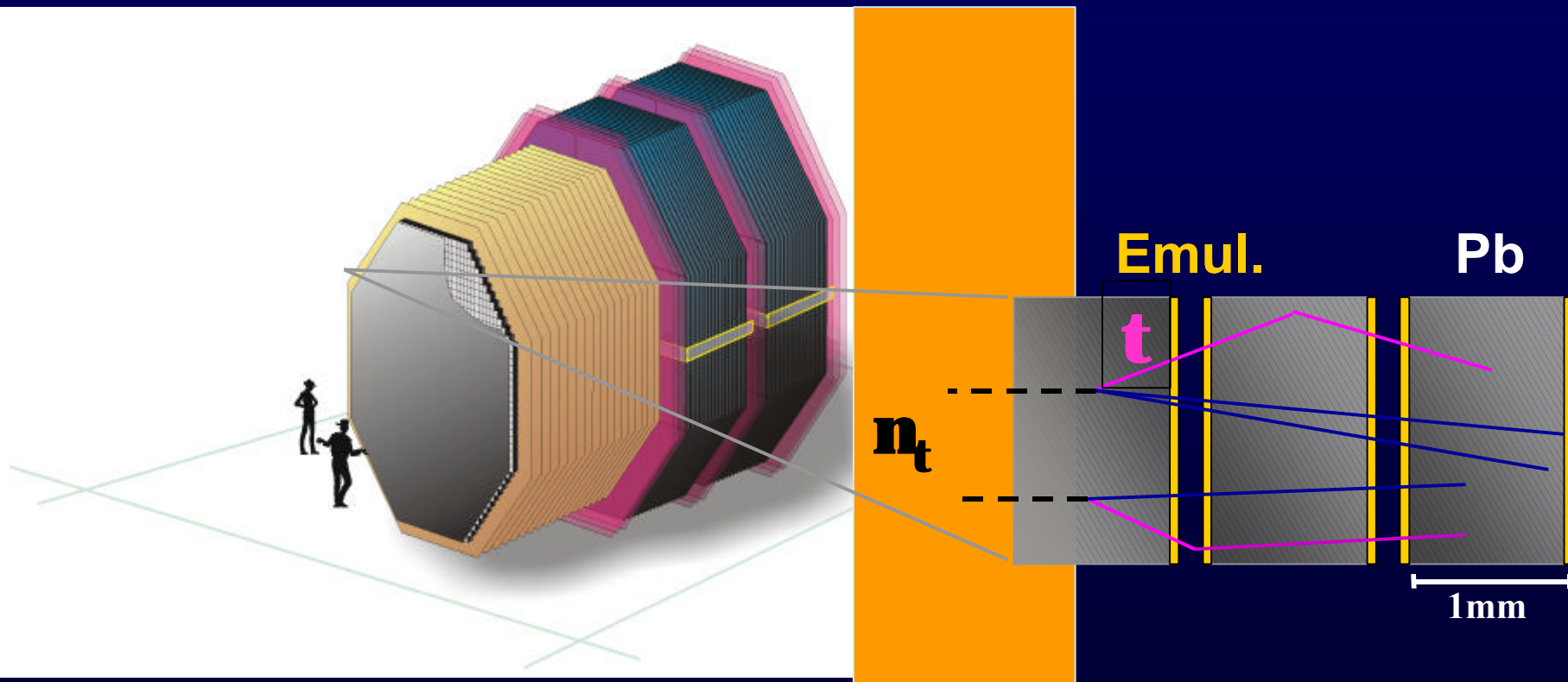
(b) very good energy and angle resolution -> ICARUS

OPERA:

walls made of bricks (total more than 200'000)

-> bricks made of "sandwiches"

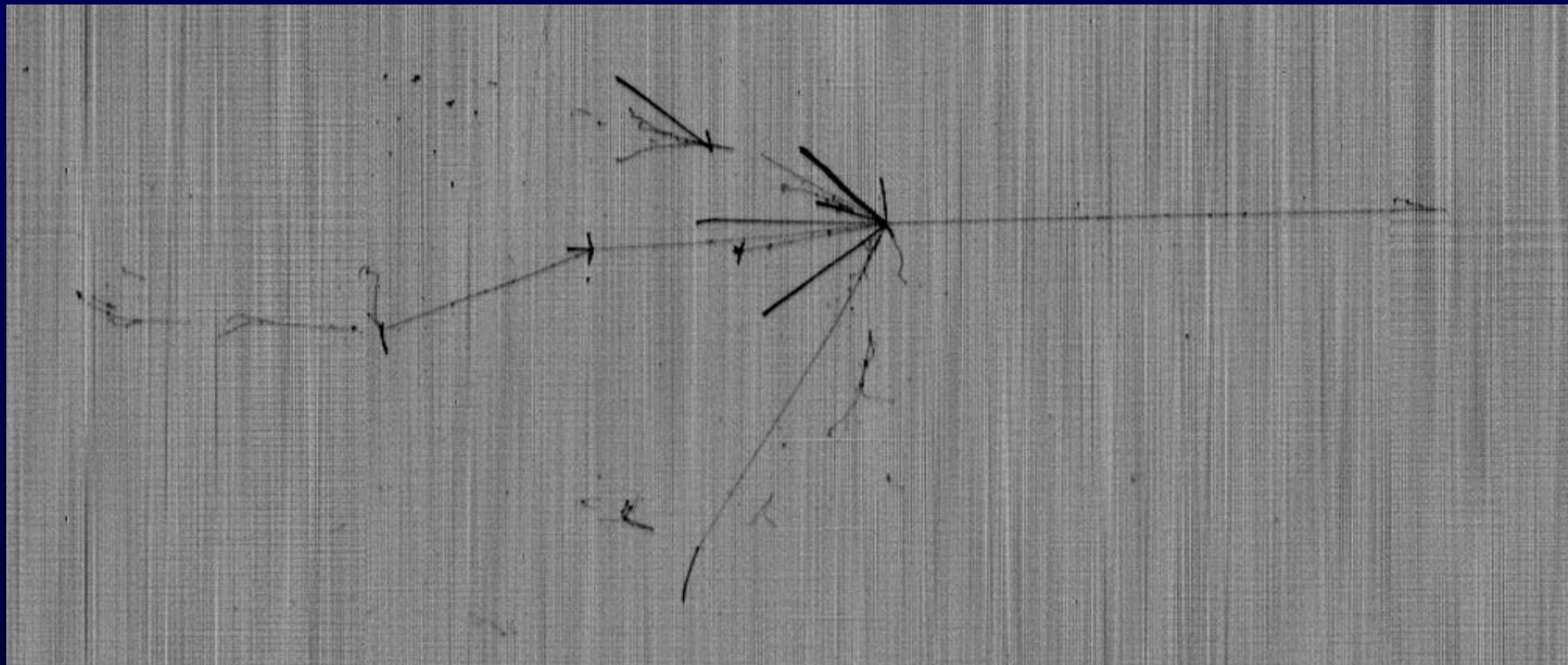
-> sandwiches made of lead and nuclear emulsion
(type of "photographic" film)



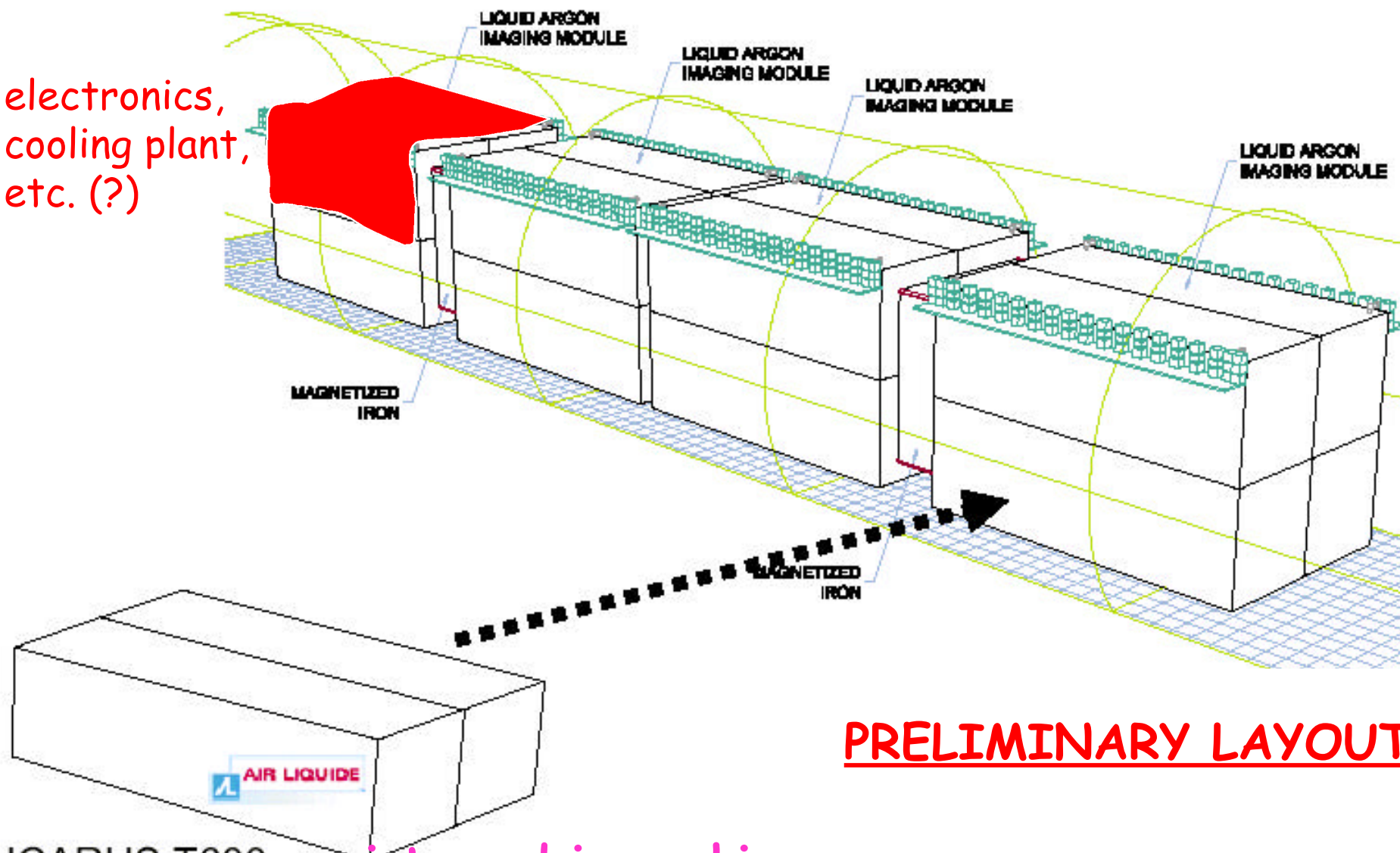
ICARUS: multi-purpose detector !



- 5000 tons ultra-pure liquid argon
- provides "electronic" picture of interactions
- > example from 600 + module (2001 - cosmic ray):



electronics,
cooling plant,
etc. (?)



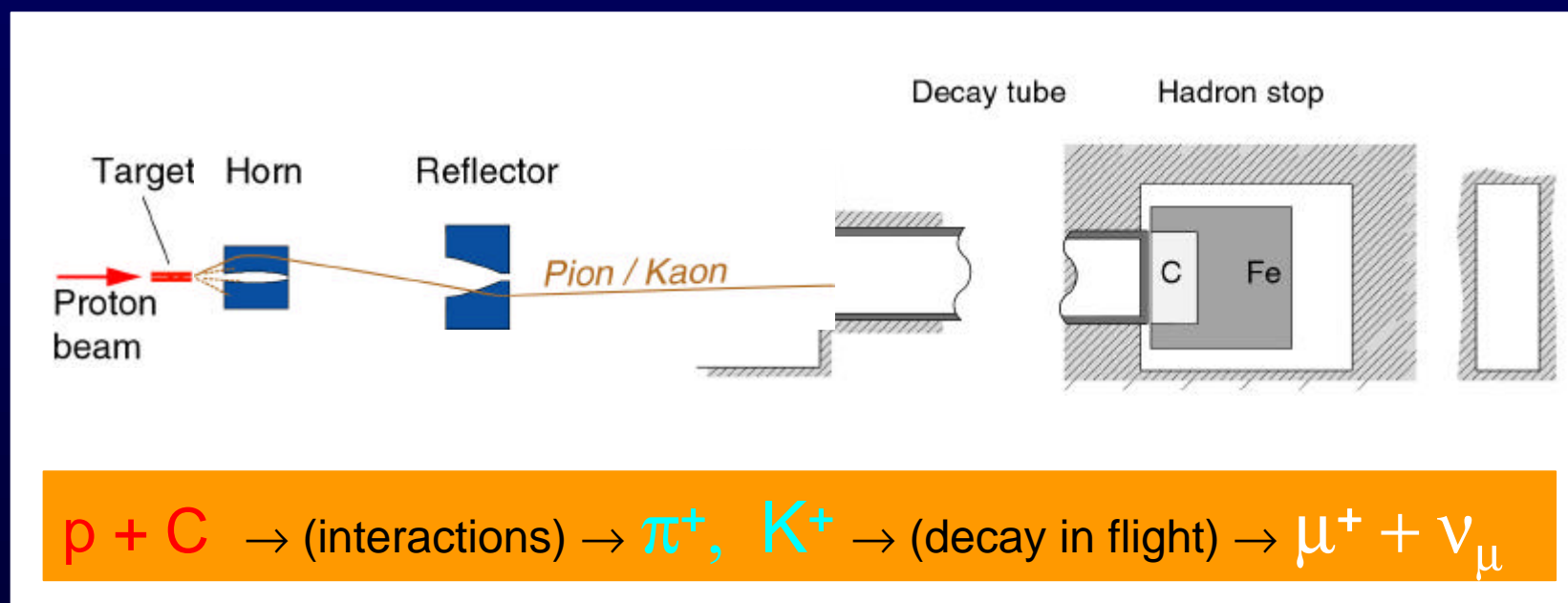
PRELIMINARY LAYOUT

ICARUS T600 - exists and is working

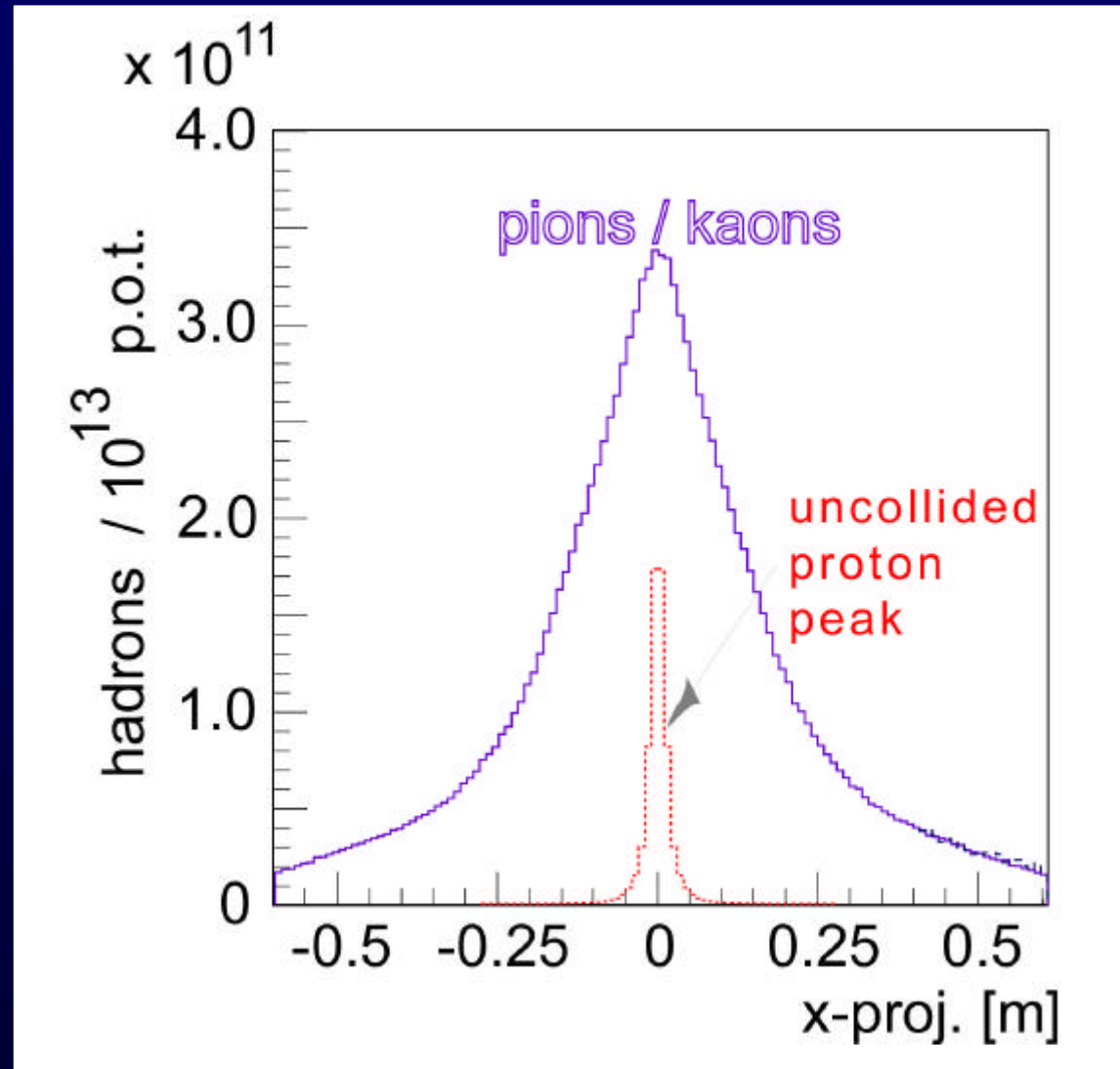
"back to CERN"...

CNGS beam-line: the main components (1)

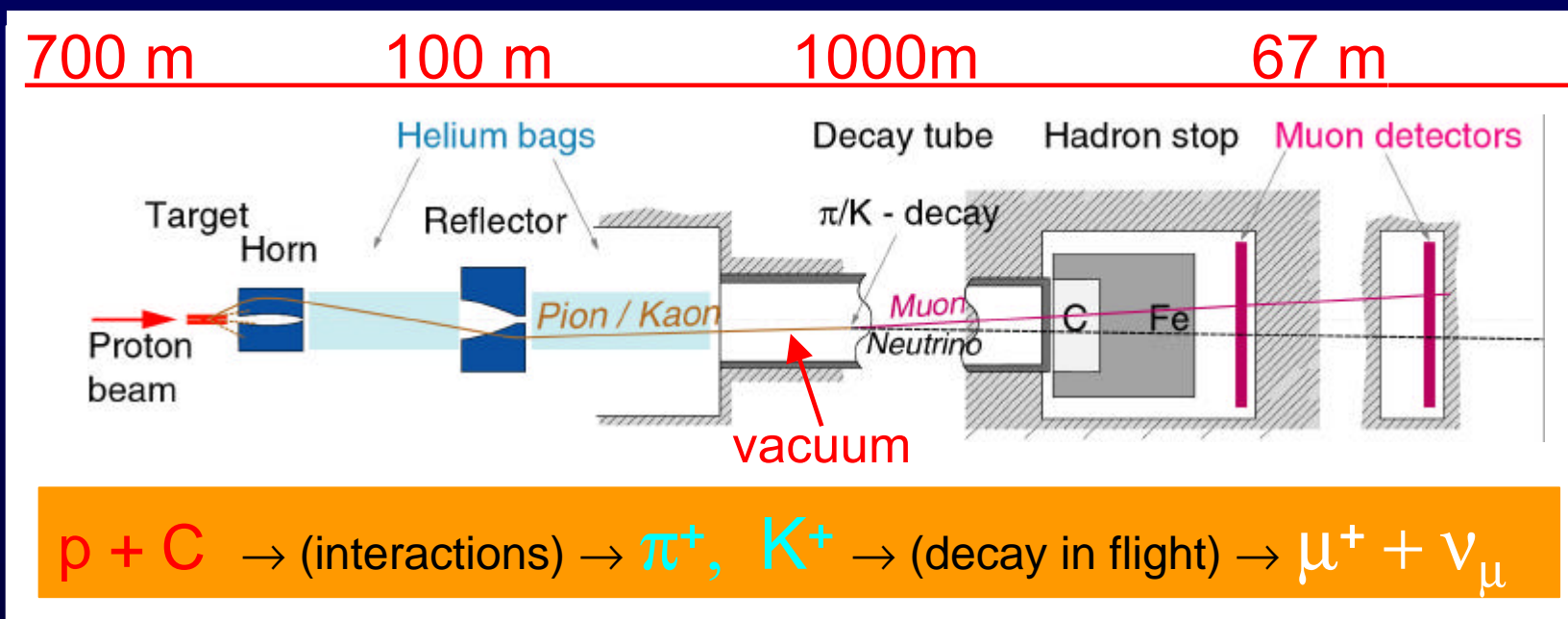
(based on CERN experience: PS / SPS neutrino beams \rightarrow WANF)



p / K profile at entrance to decay tunnel

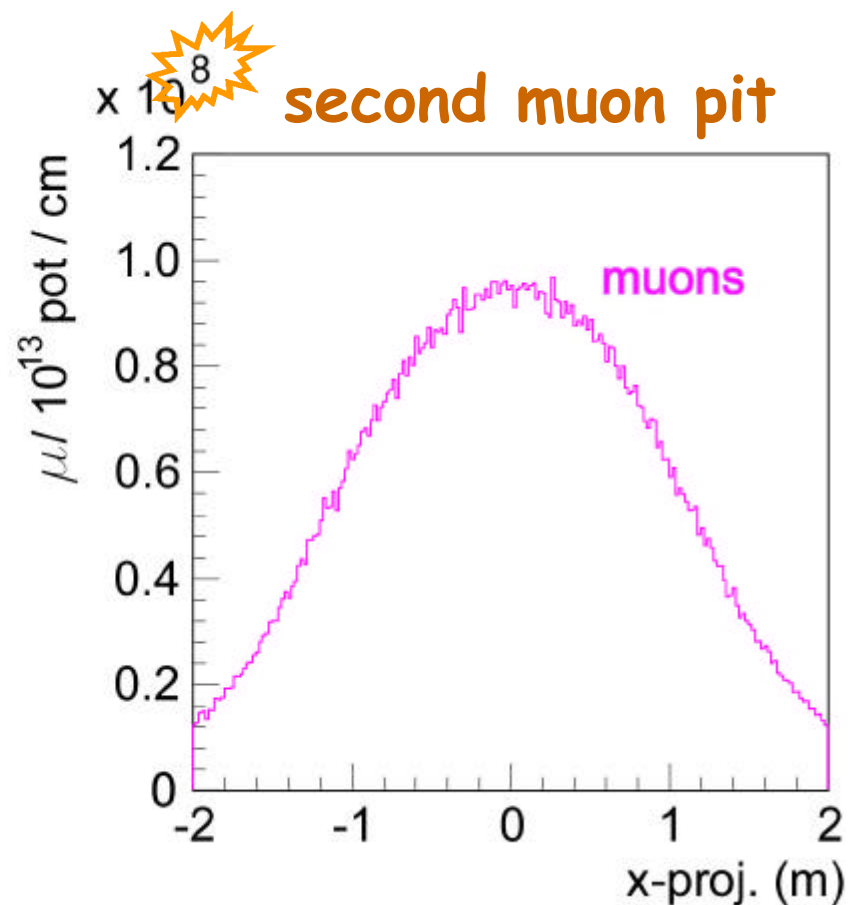
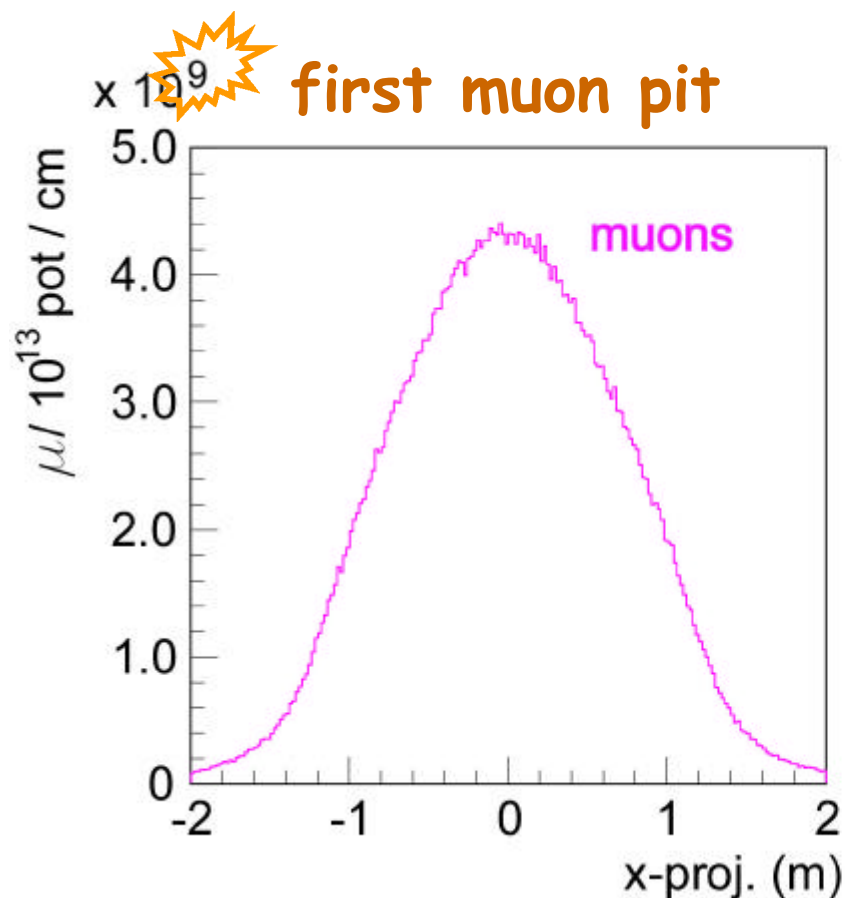


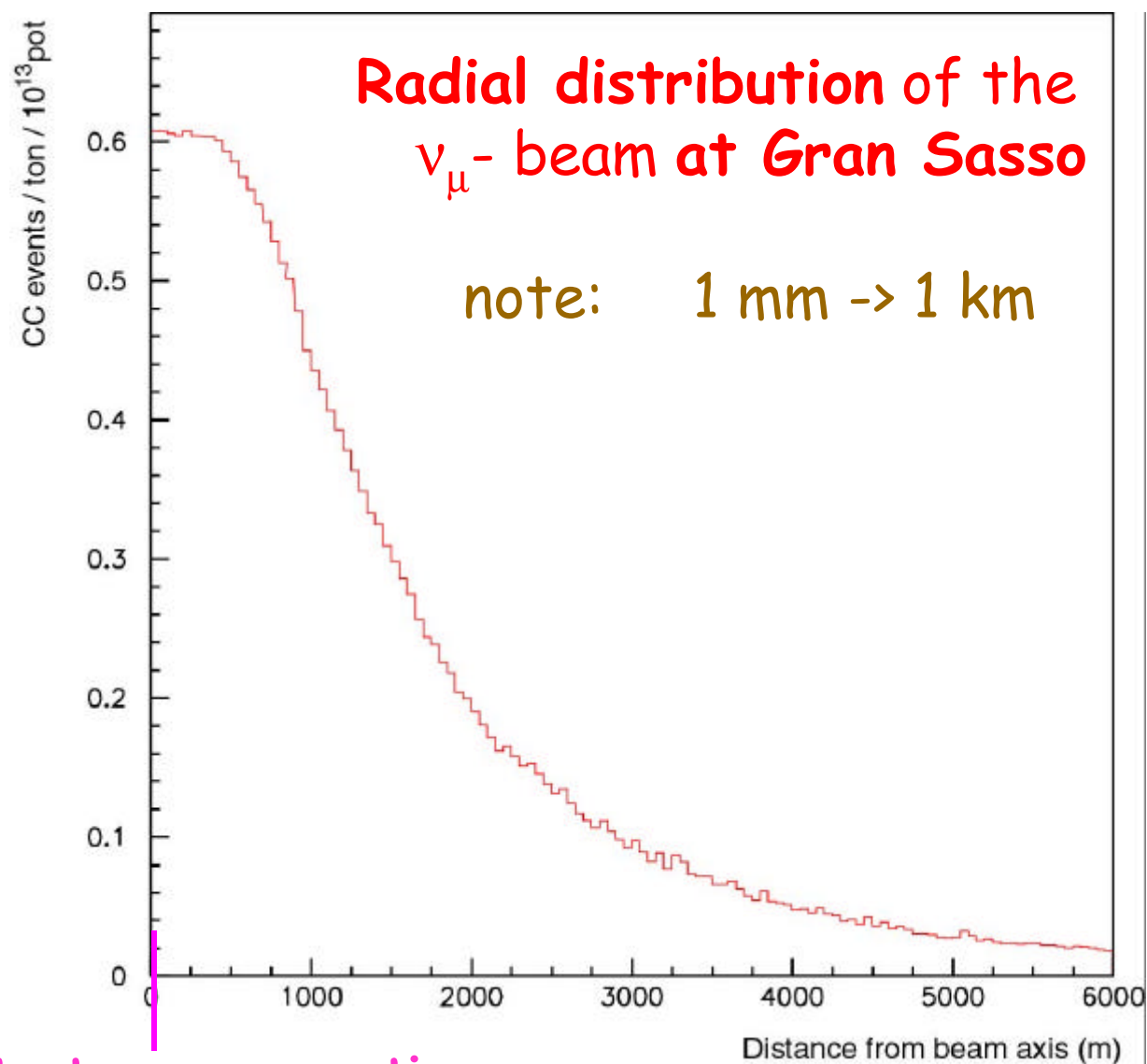
CNGS: the main components (2)



$$\pi^+ \rightarrow \mu^+ + \nu_\mu$$

expected CNGS muon profiles





detector cross section

Number of particles expected per year:



For 1 year of CNGS operation, we expect:

(4.8×10^{13} protons in SPS, 55% efficiency -- 1997)

protons on target

4.5×10^{19}

pions / kaons at entrance to decay tunnel

5.8×10^{19}

muons in first / second muon pit

$3.6 \times 10^{18} / 1.1 \times 10^{17}$

ν_{μ} in 100 m^2 at Gran Sasso

3.5×10^{12}

ν_{μ} "charged current" **events** per 1000 +
($\nu + N \rightarrow N' + \mu$)

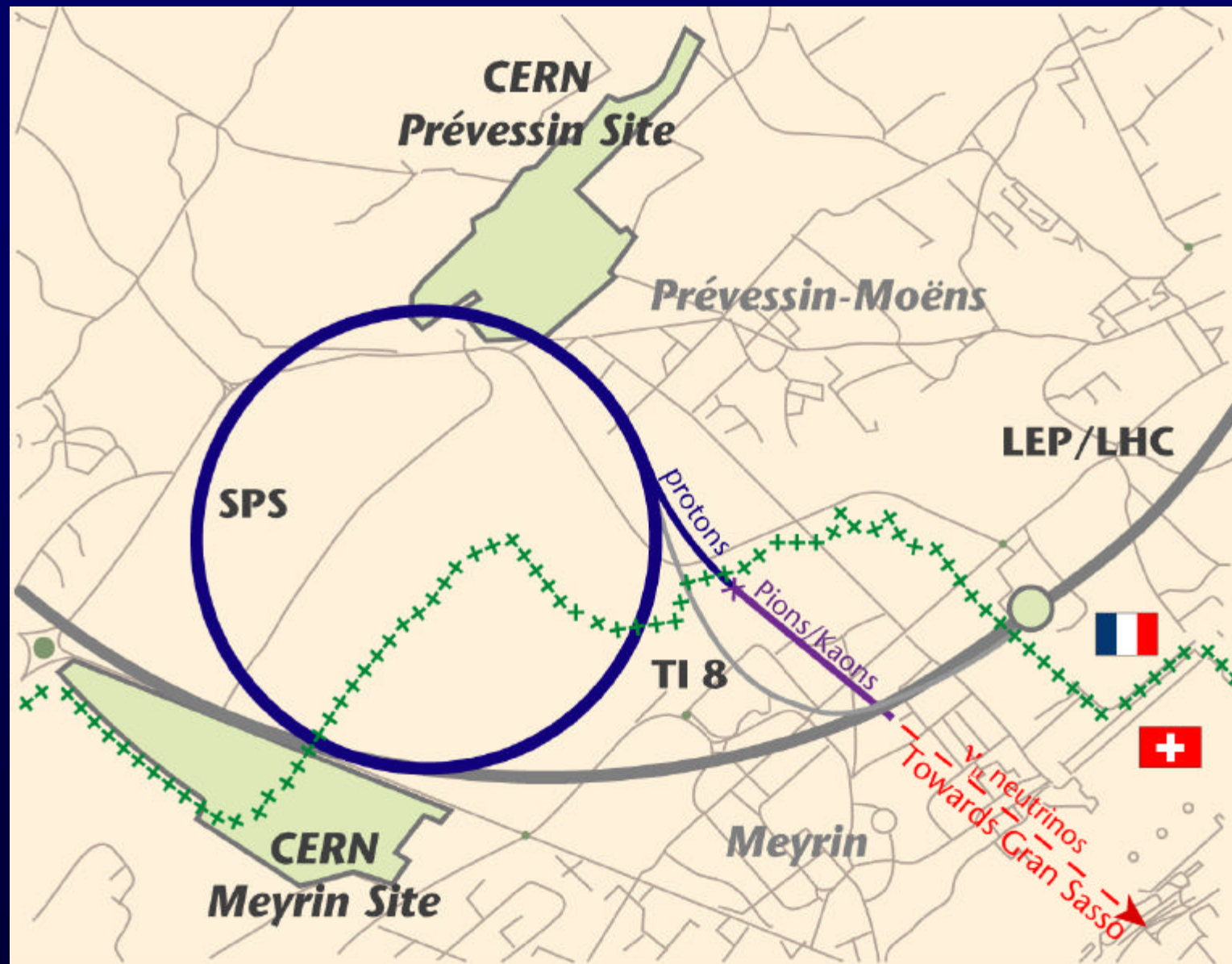
≈ 2500

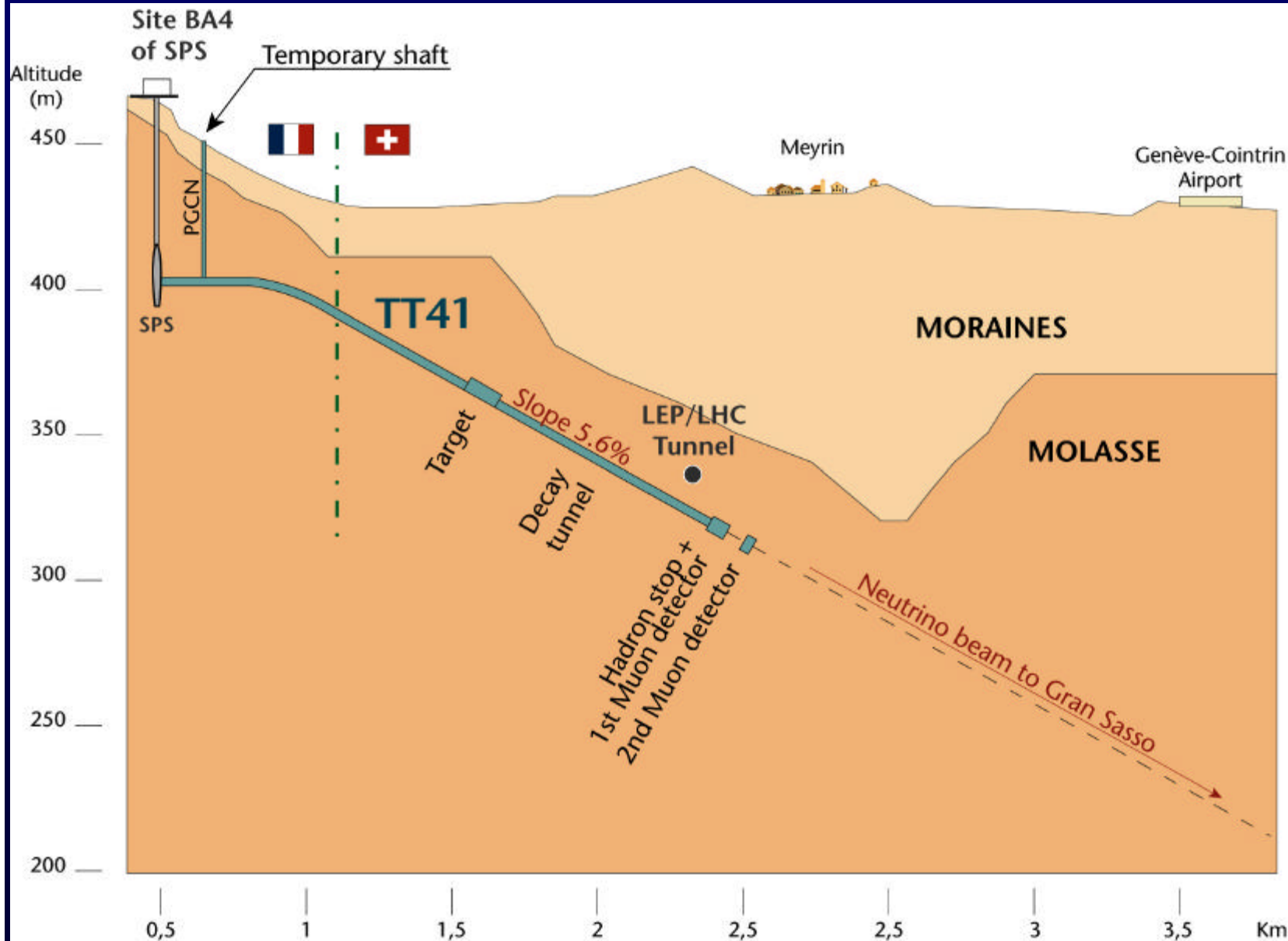
ν_{τ} events (from oscillation)

≈ 20 "detectable"

n_t events detected "in OPERA"

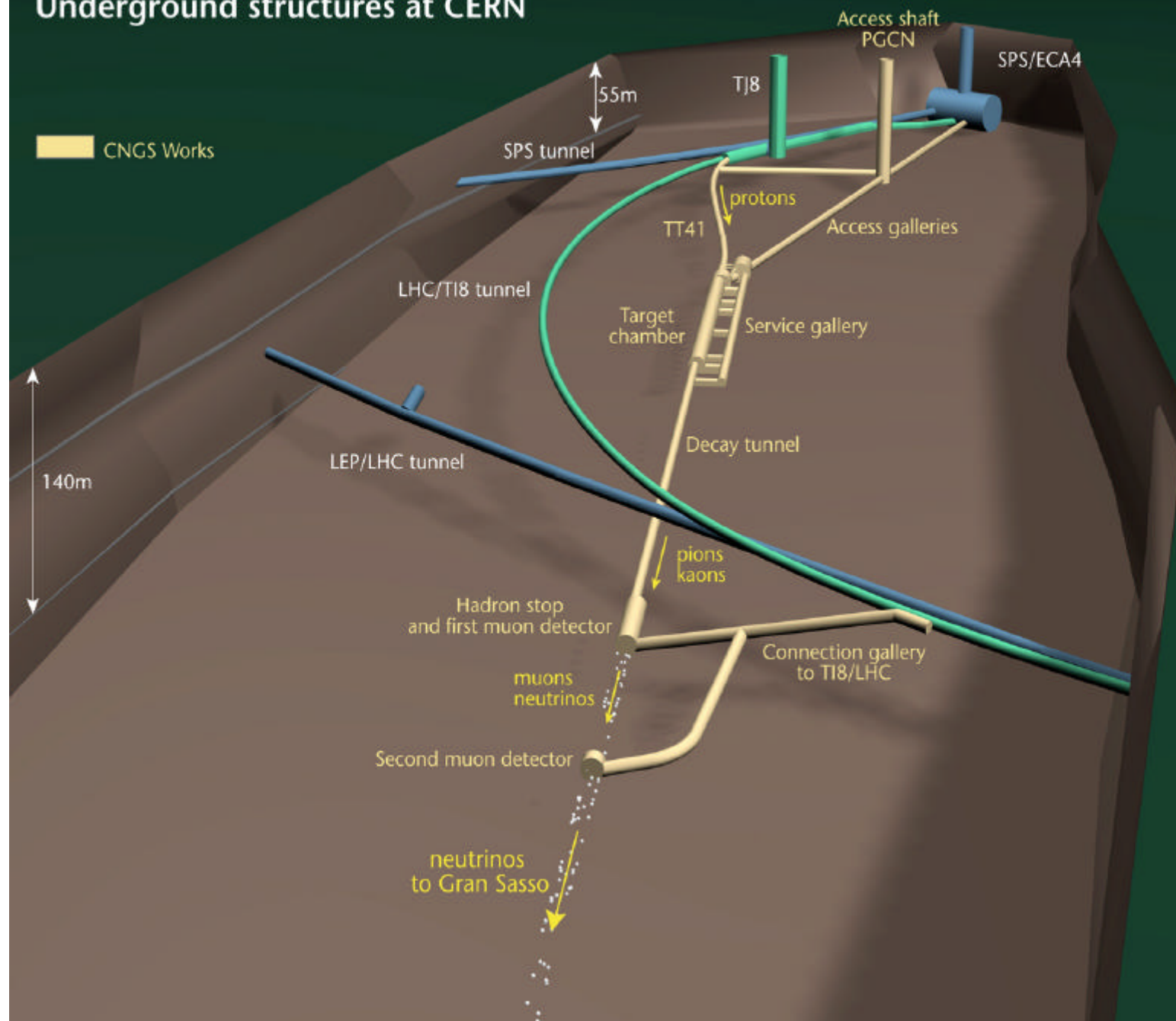
≈ 2.5 (b.g. 0.15)





CERN NEUTRINOS TO GRAN SASSO

Underground structures at CERN



CNGS civil engineering

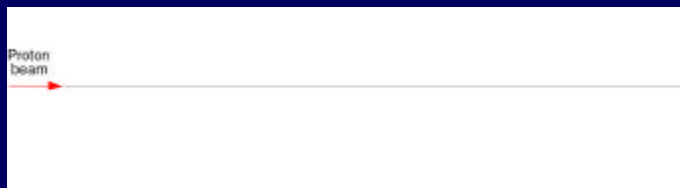
more than 3 km of tunnels and caverns
(diameter 3.1 m \rightarrow 6.0 m)

more than 45'000 m³ rock
to be removed

more than 12'000 m³ concrete
to be "sprayed" or poured



CNGS -- proton beam (1)



- > protons from SPS
- > **NEW** fast extraction **under construction in point 4**
- > this extraction system **needed for TI8** -> LHC transfer line (but modified for CNGS)
- > CNGS p-beam branches off after ≈ 100 metres
- > 700 metres of proton beam line **to CNGS target**

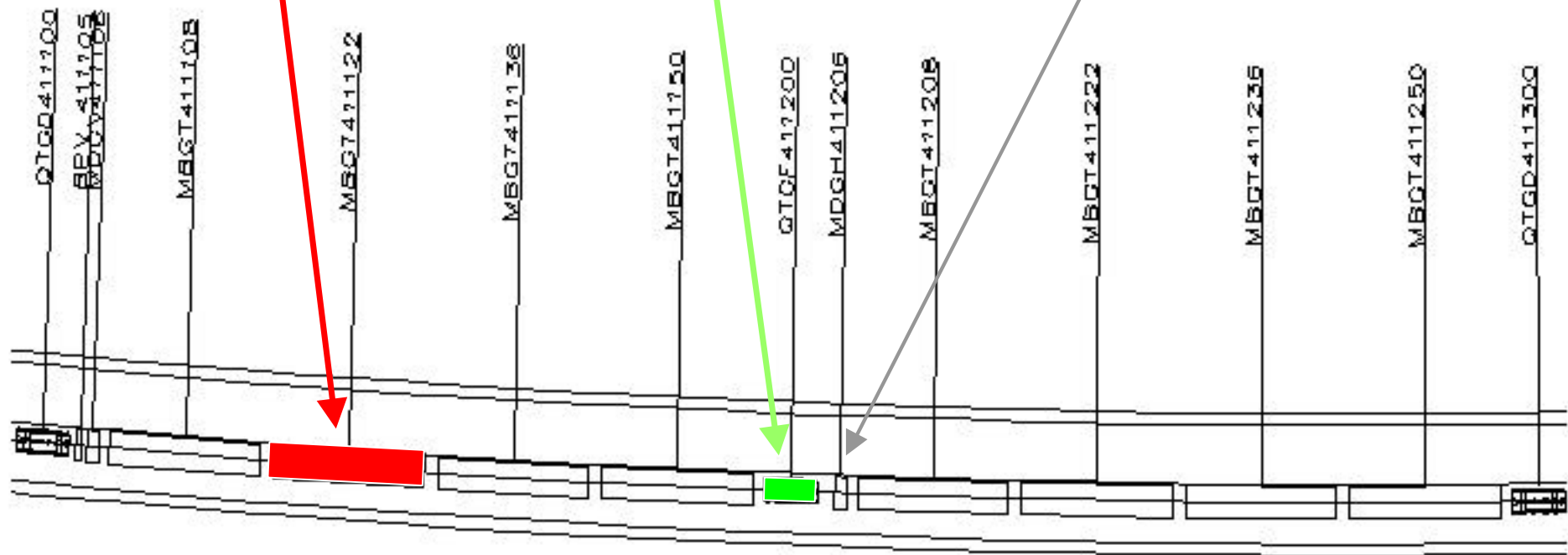
CNGS -- proton beam (2)

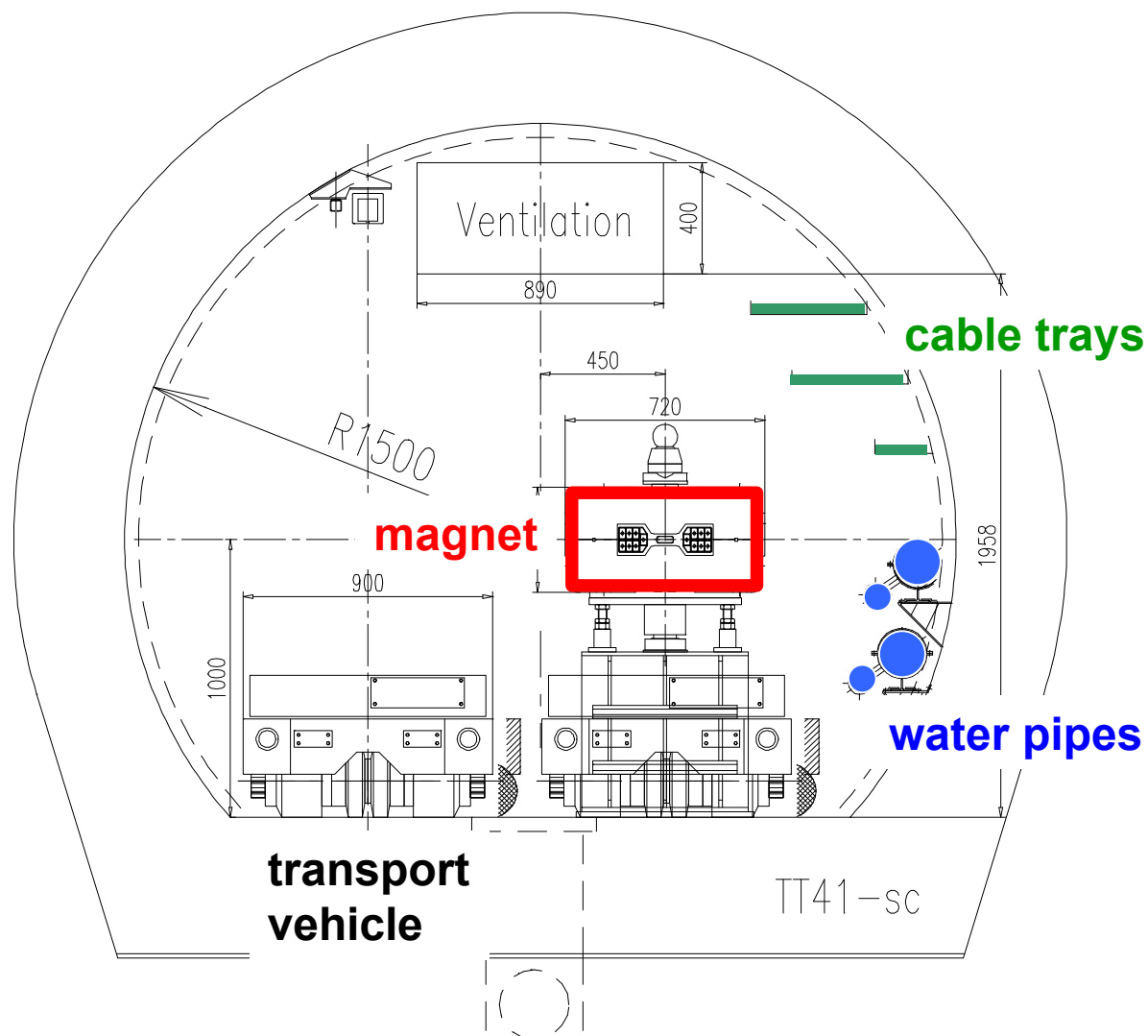


73 deflection (dipole) magnets (6.4 m long) +

21 quadrupole magnets + correction dipoles

+vacuum + beam observation + ...





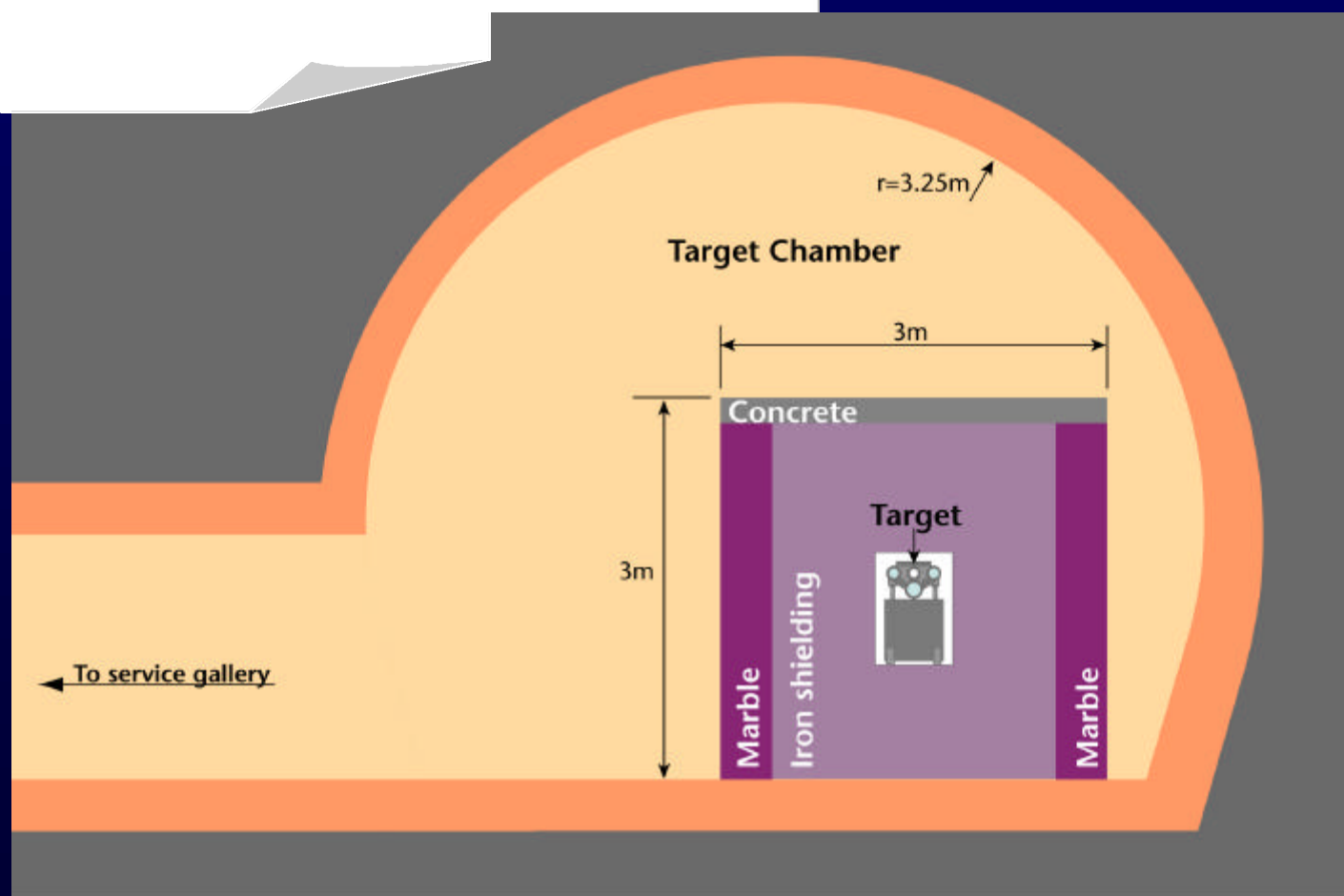
proton beam tunnel
cross section
(preliminary)

warning:
there is "no space"

CNGS -- target station (1)



Proton beam
Target



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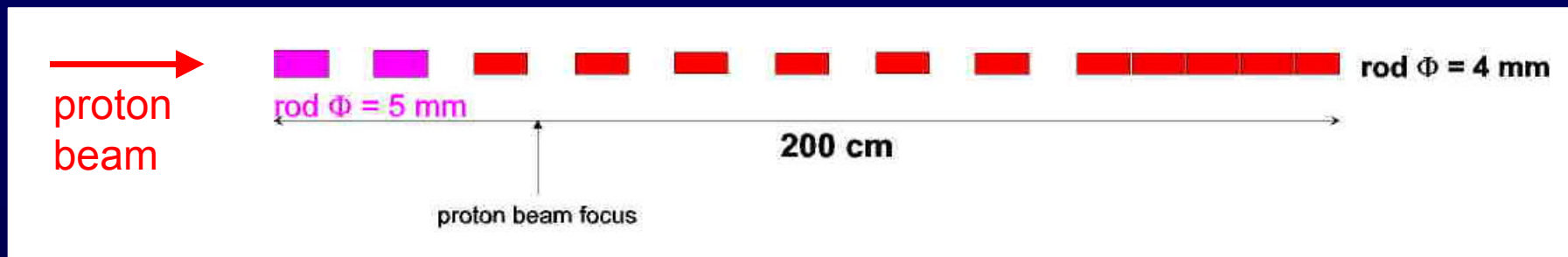
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CNGS -- target station (2)

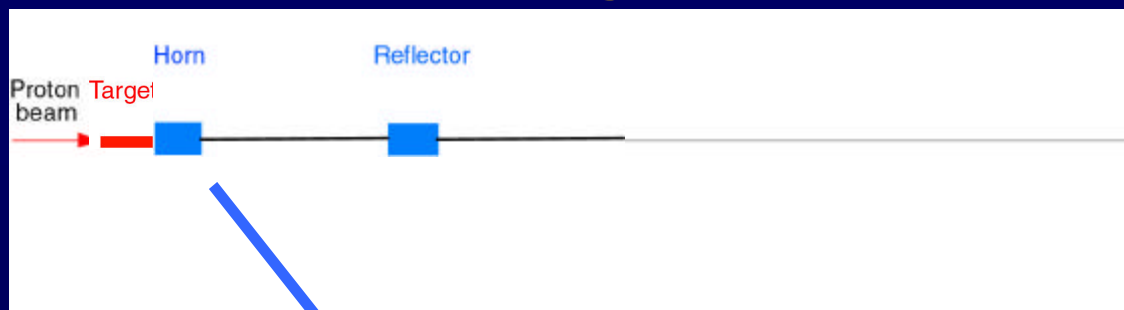


-> 10 cm long graphite rods, $\varnothing = 5\text{mm}$ and/or 4mm



Note: - target rods interspaced to "let the pions out"
- target is helium cooled
(remove heat deposited by the particles)

CNGS -- focusing devices (1)

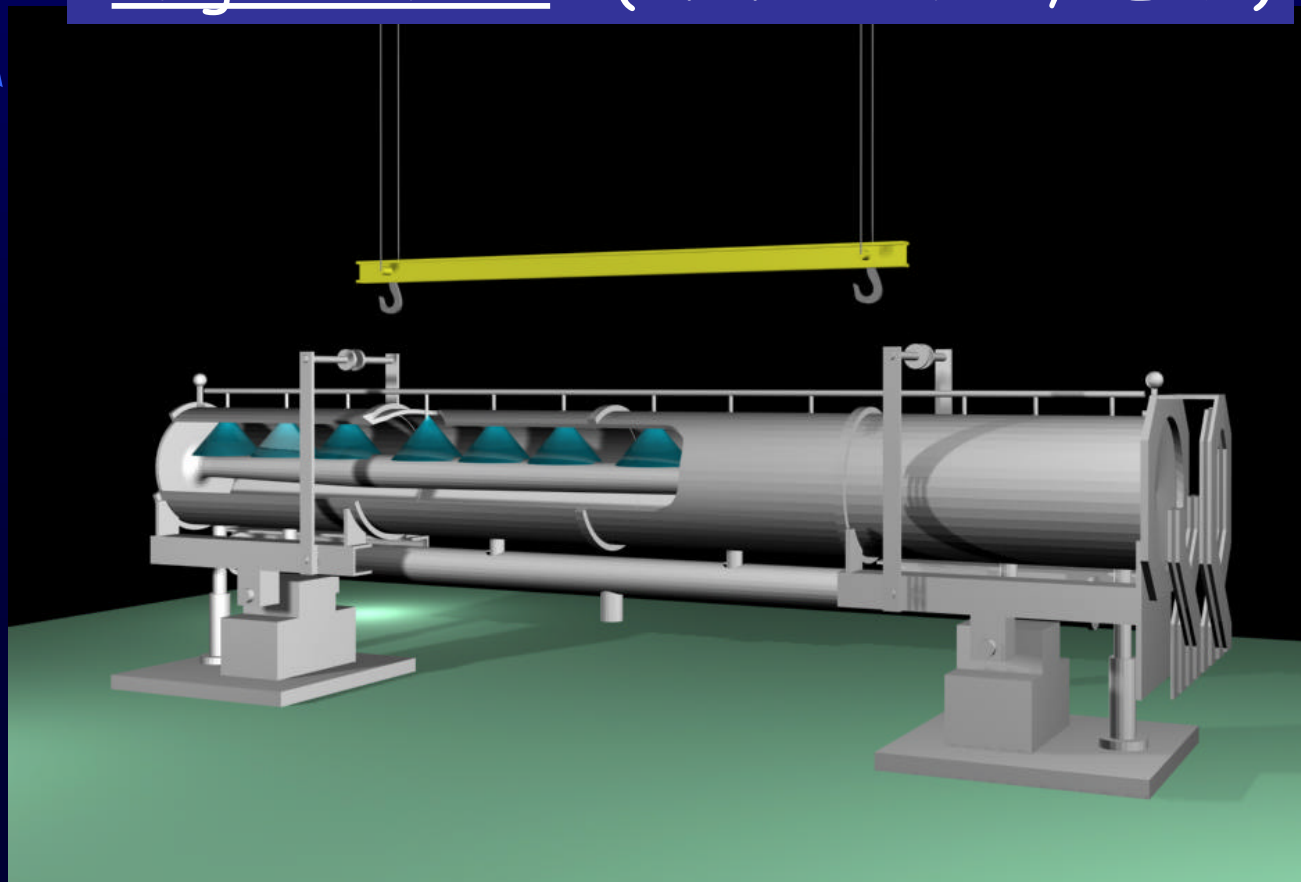


"Magnetic Horn" (S. v.der Meer, CERN)

length: 6.5 m
diameter: 70 cm
weight: **1500 kg**

Pulsed devices:
150kA / 180 kA, 1 ms

water-cooled:
distributed nozzles



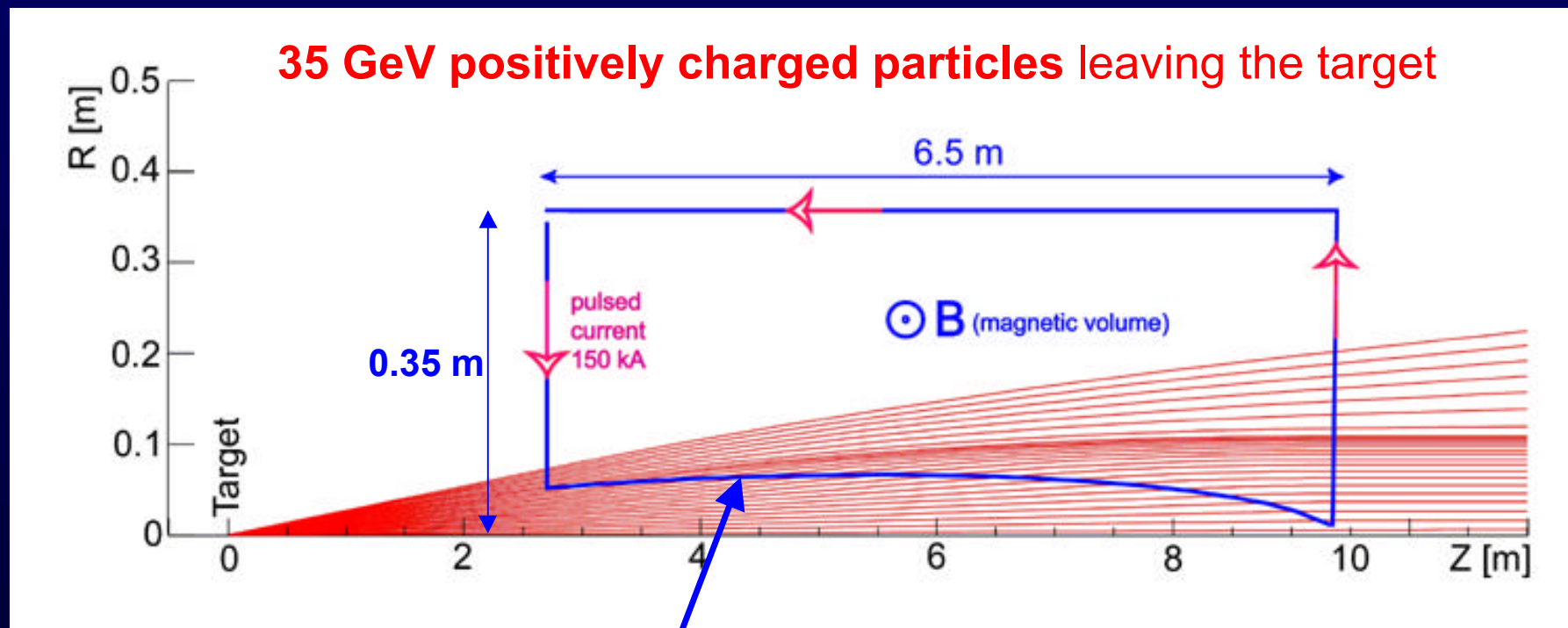
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Principle of focusing with a Magnetic Horn



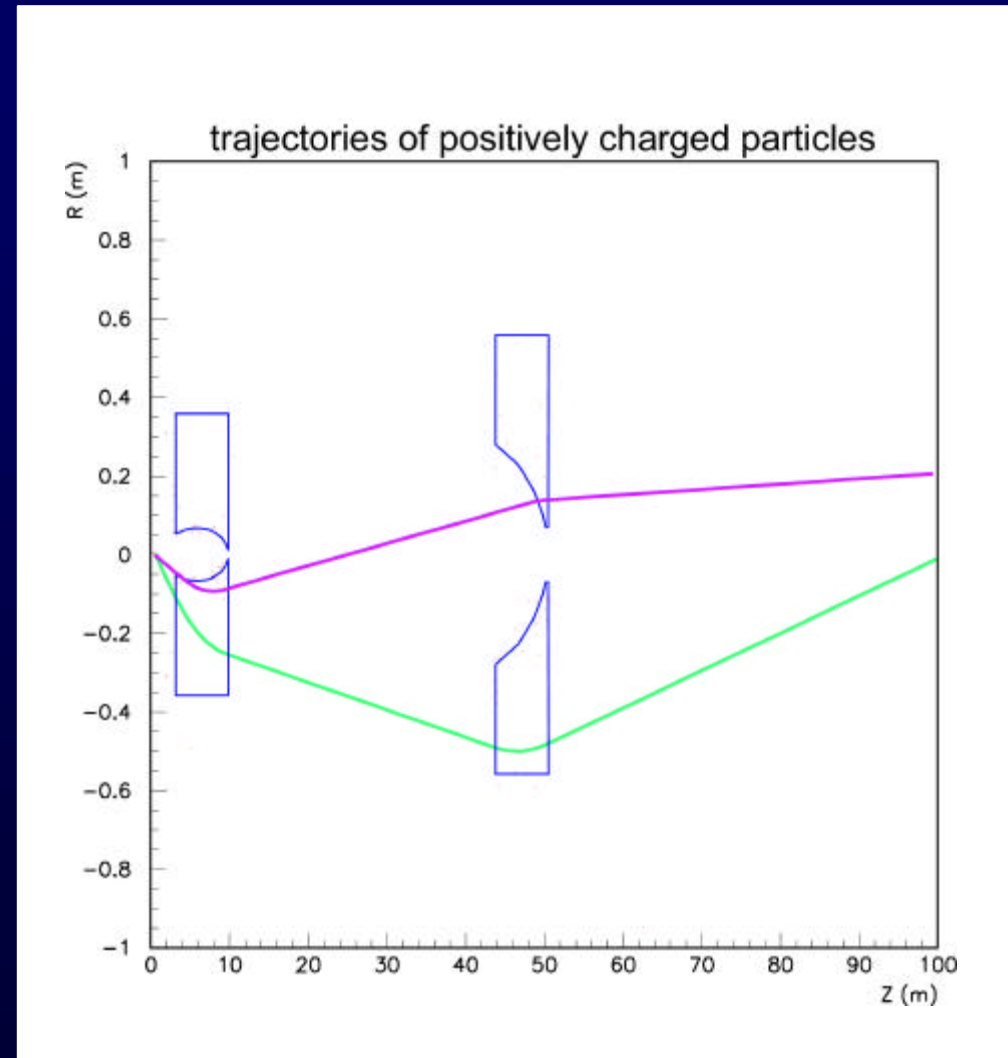
magnetic volume given by "one turn" at high current:

- ♦ specially shaped inner conductor
 - end plates
- ♦ cylindrical outer conductor

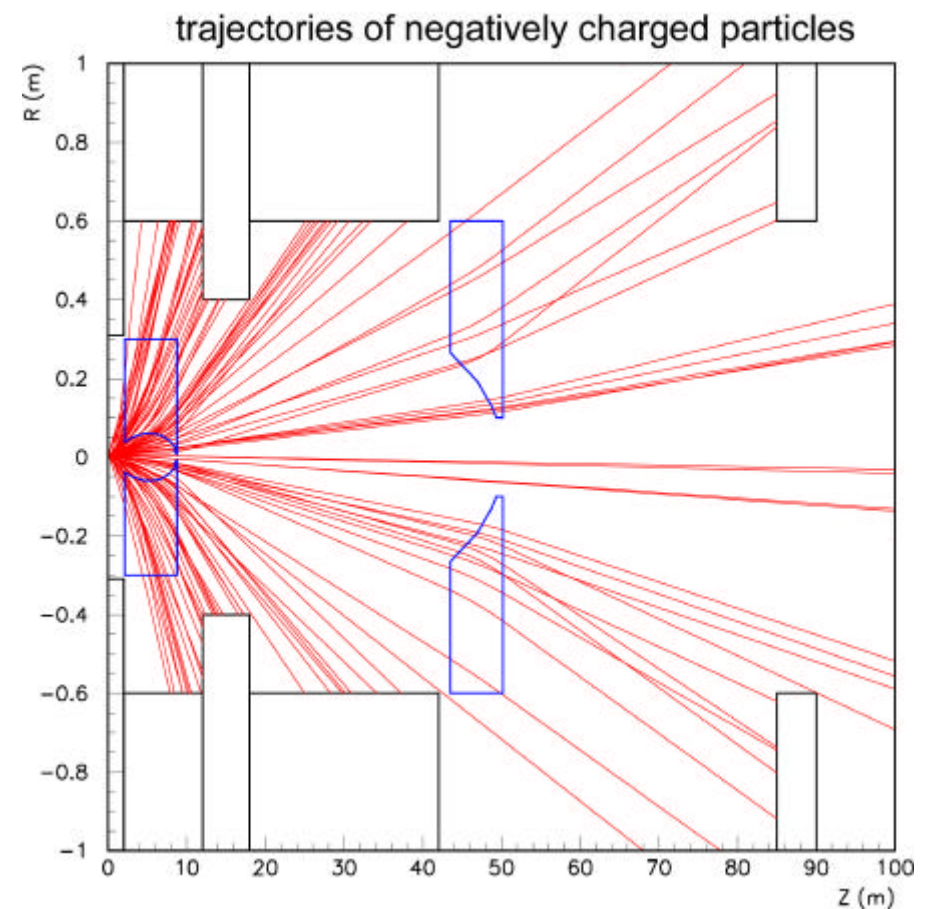
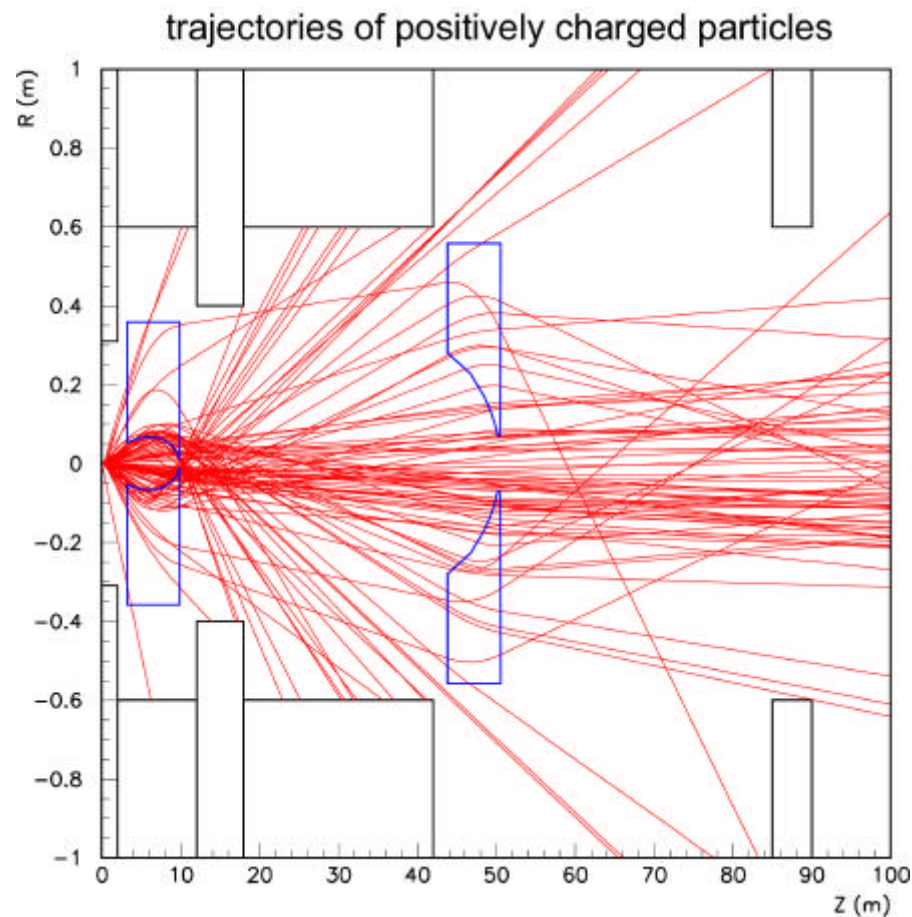


inner conductor

secondary beam focusing with horn/reflector



secondary beam focusing with horn/reflector

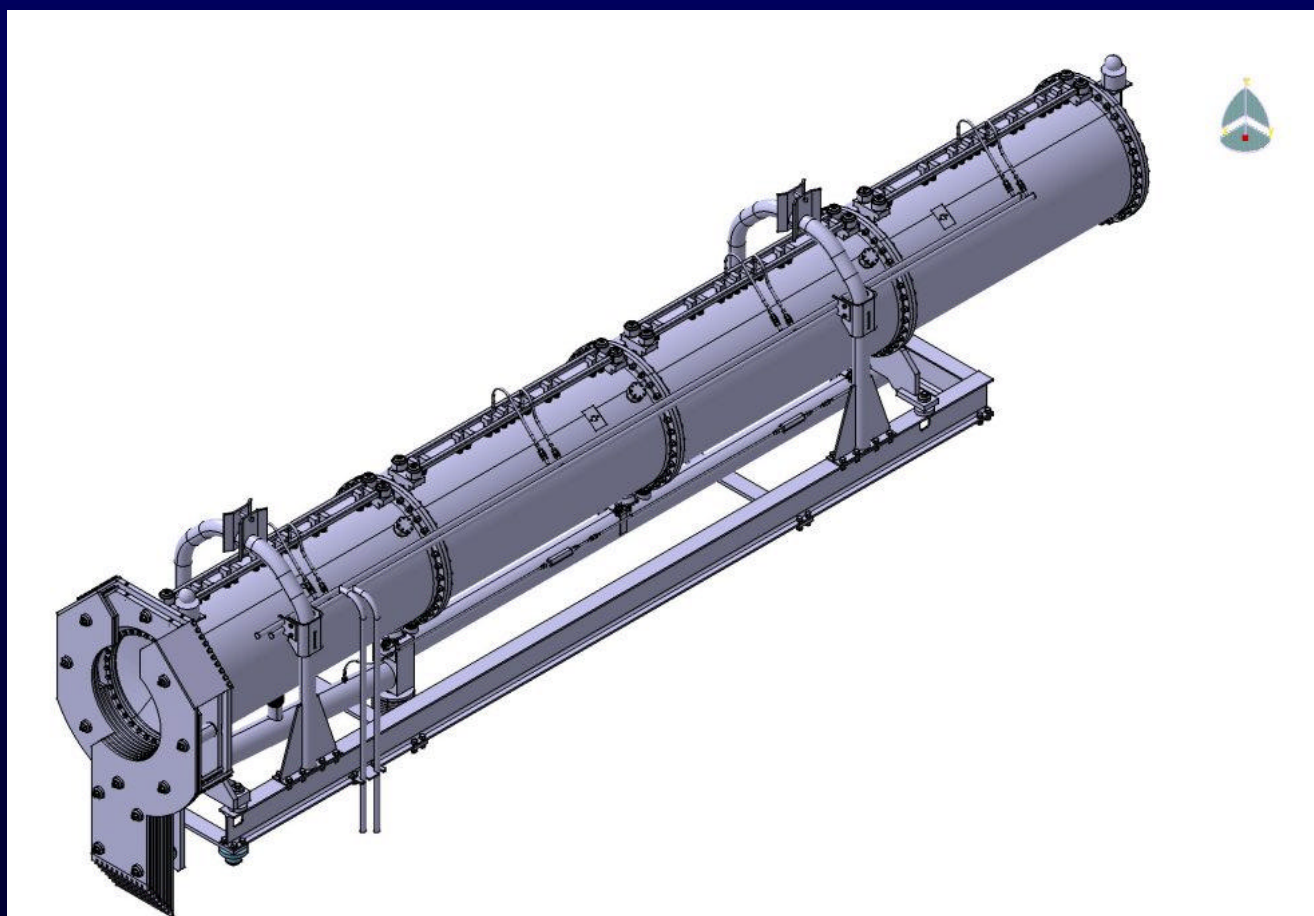


CNGS -- focusing devices (2)

(collaboration with IN2P3, Paris)

Design criteria:

>95% probability to work for 5×10^7 pulses

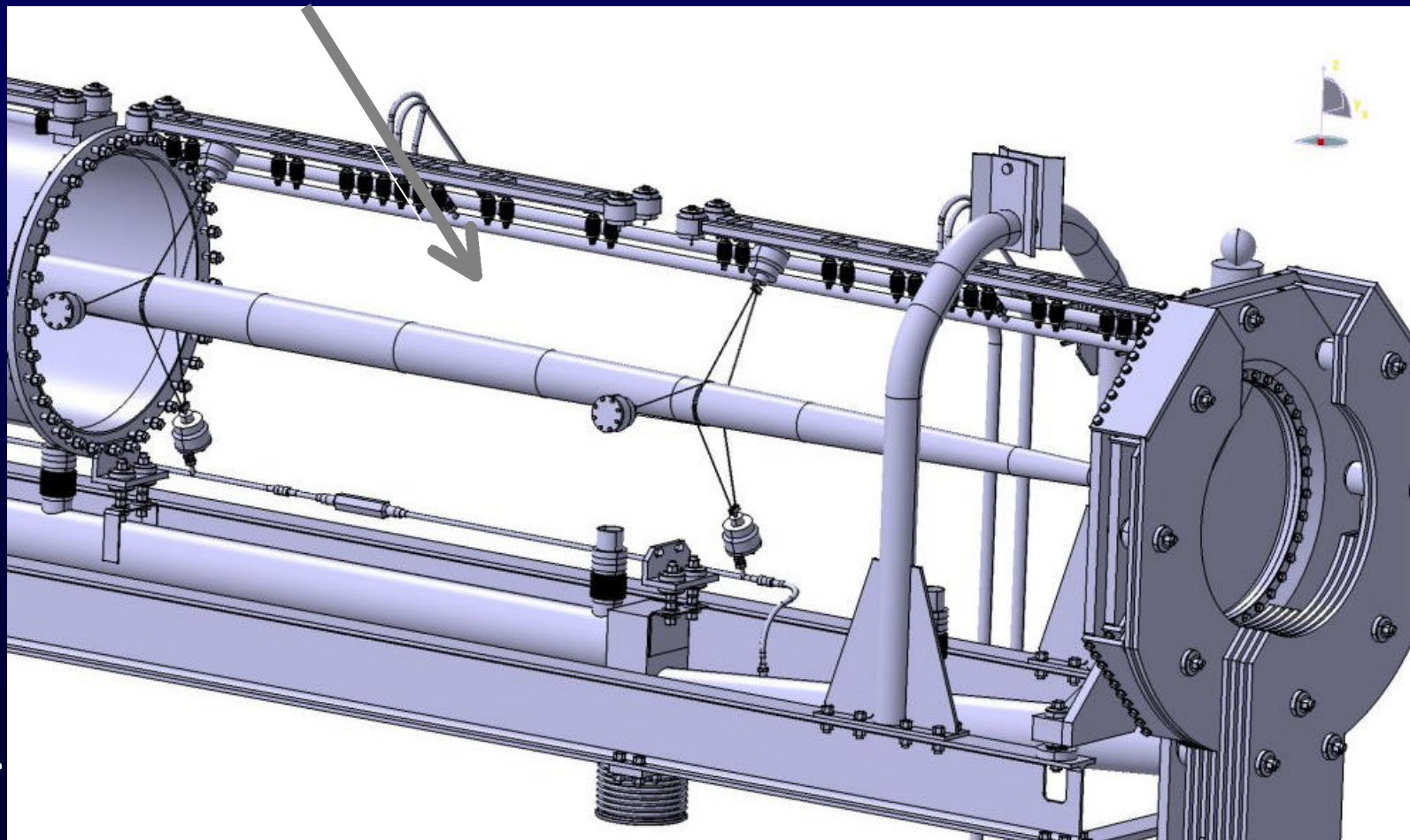


CNGS -- focusing devices (2)

(collaboration with IN2P3, Paris)

The inner conductor:

- as thin as possible (particle absorption)
- as thick as necessary (mechanical stability)

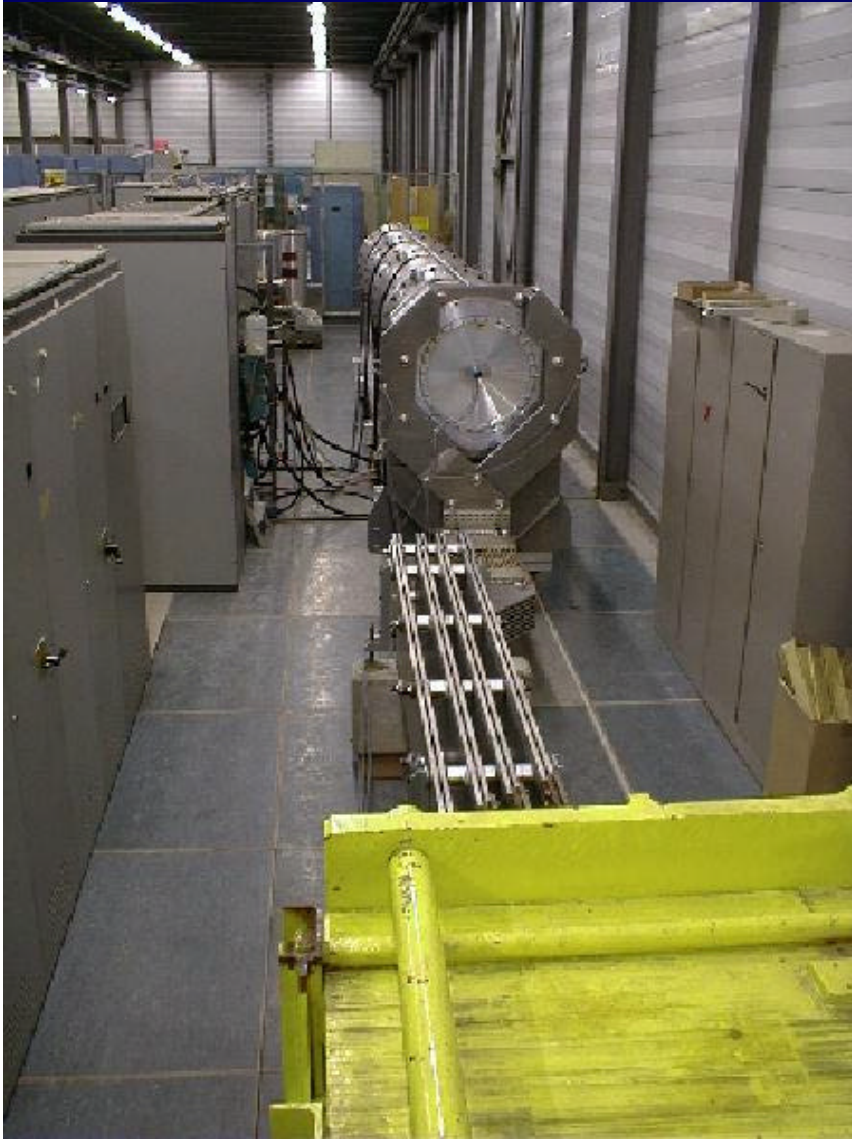


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CNGS -- focusing devices (3)



Horn prototype tests in BA7:
1.5 Mio pulses in 2000

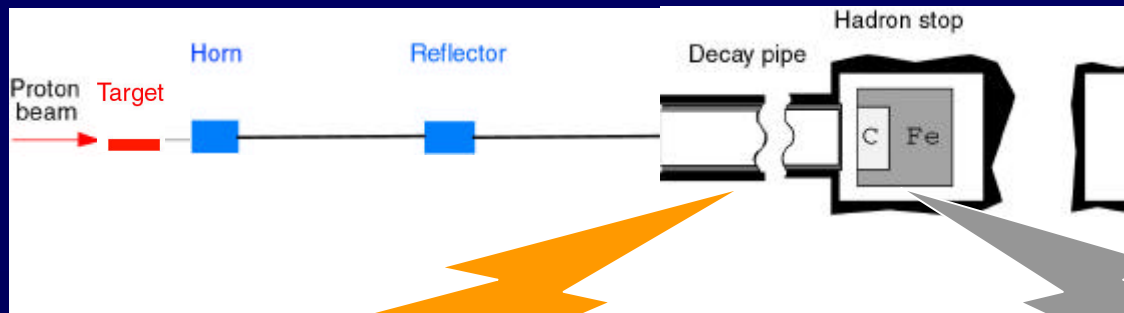


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CNGS -- decay tube + hadron stop



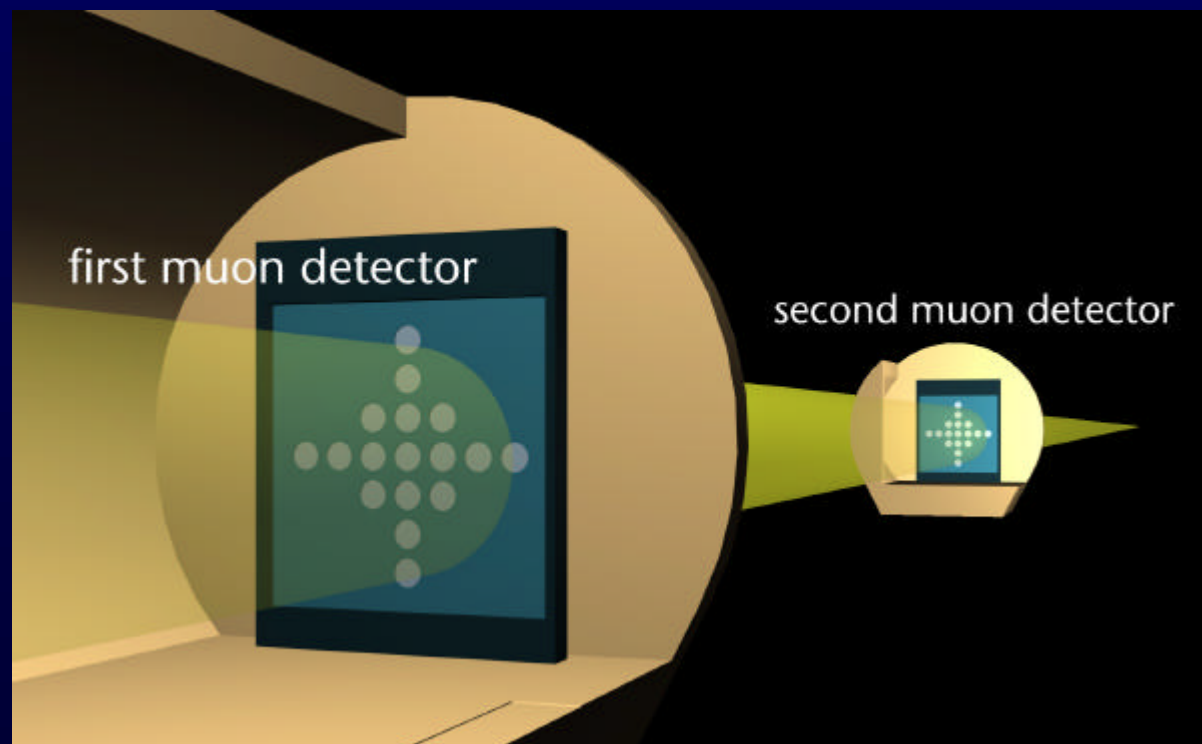
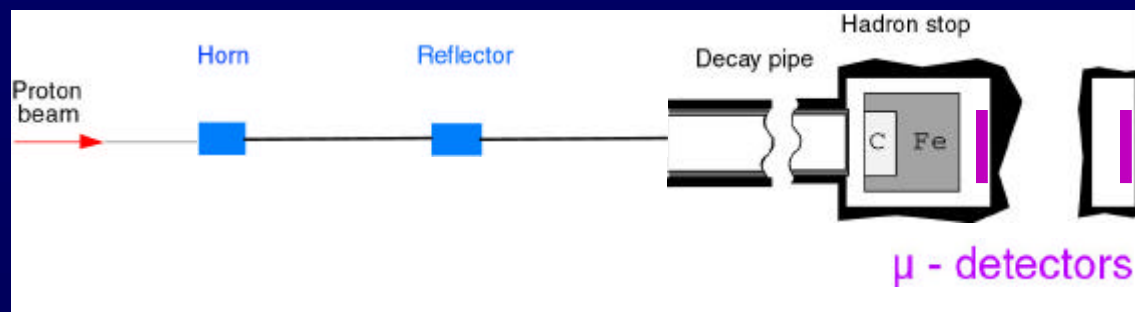
- dimensions of decay tube:

- 2.45 m diameter steel tubes, 6 m long pieces, **1 km total**
- welded together in-situ
- vacuum: ~1 mbar
- tube embedded in concrete

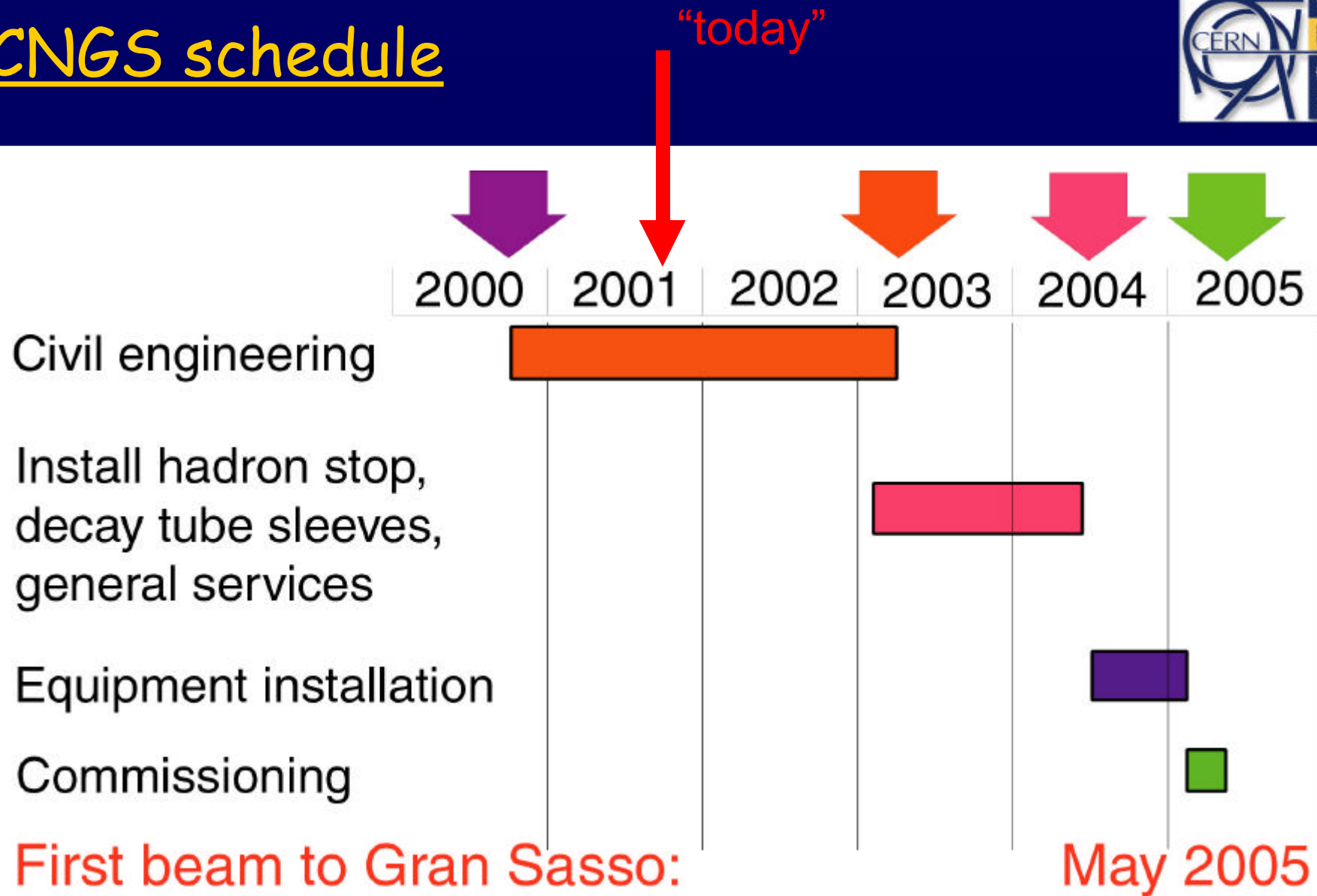
- hadron stop:

- 3.2 m graphite
- 15 m iron blocks
- upstream end: water cooled

CNGS -- muon detectors



CNGS schedule



CNGS status -- Civil Engineering (1)



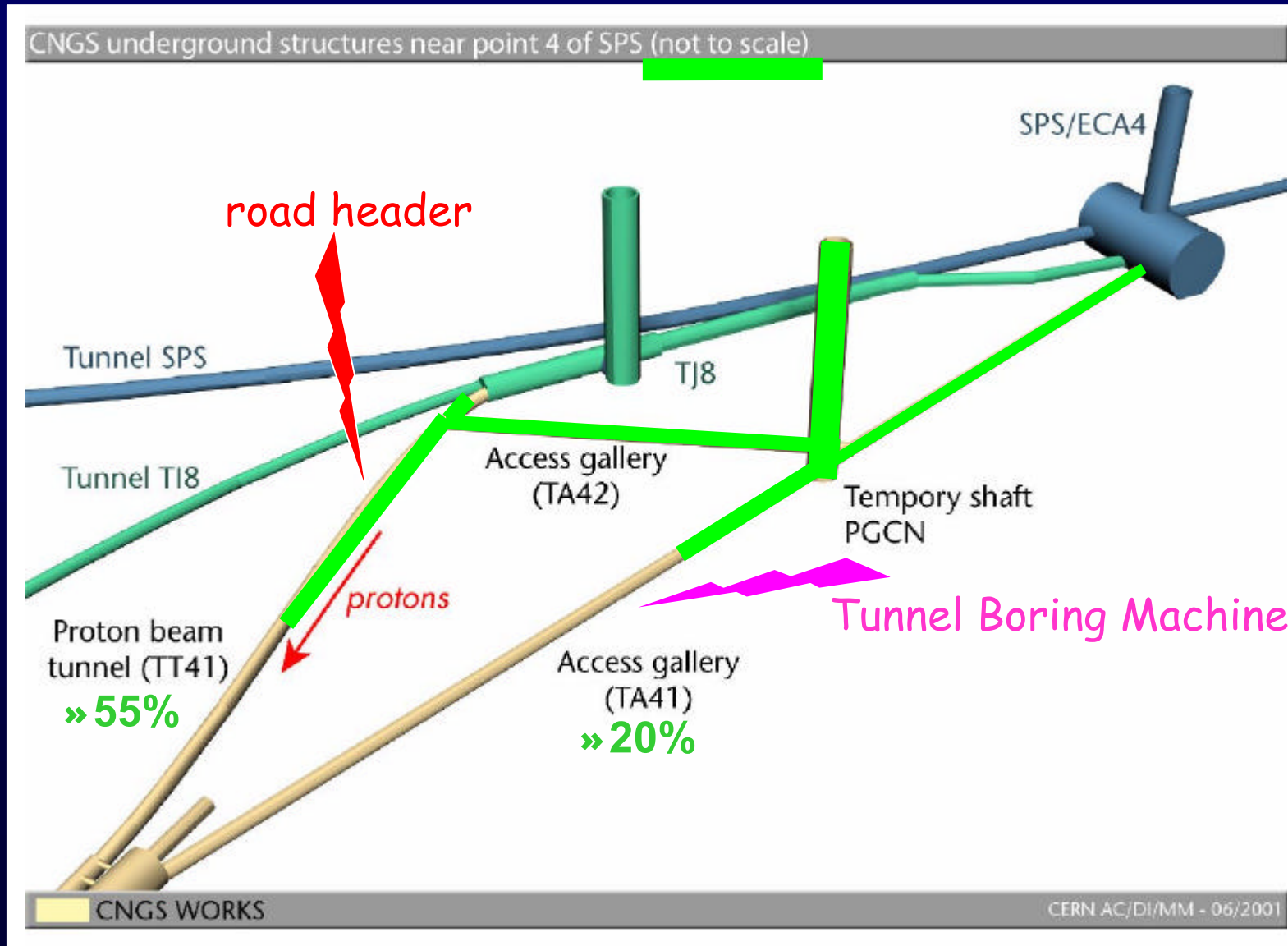
- ◆ ground breaking ceremony:
12 October 2000



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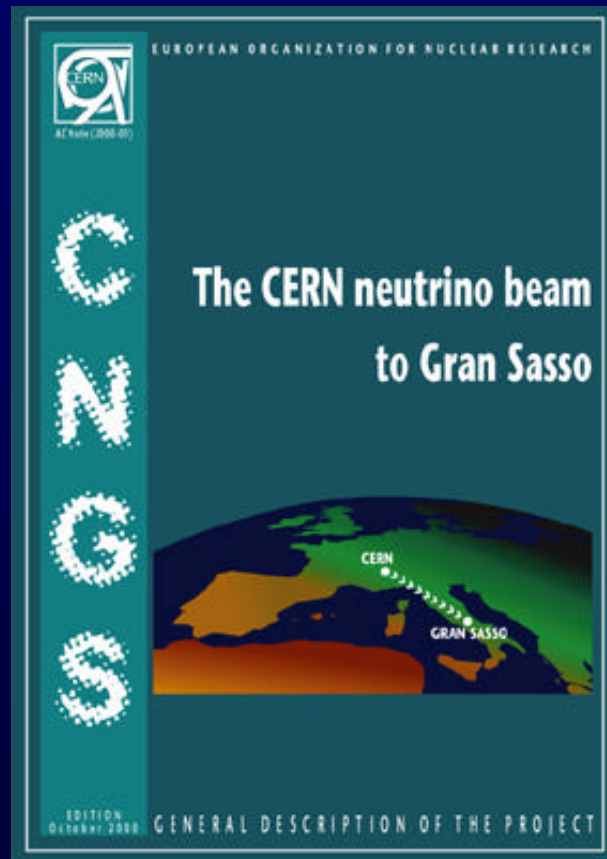
CNGS status -- Civil Engineering (2)



CNGS status -- Civil Engineering (3)

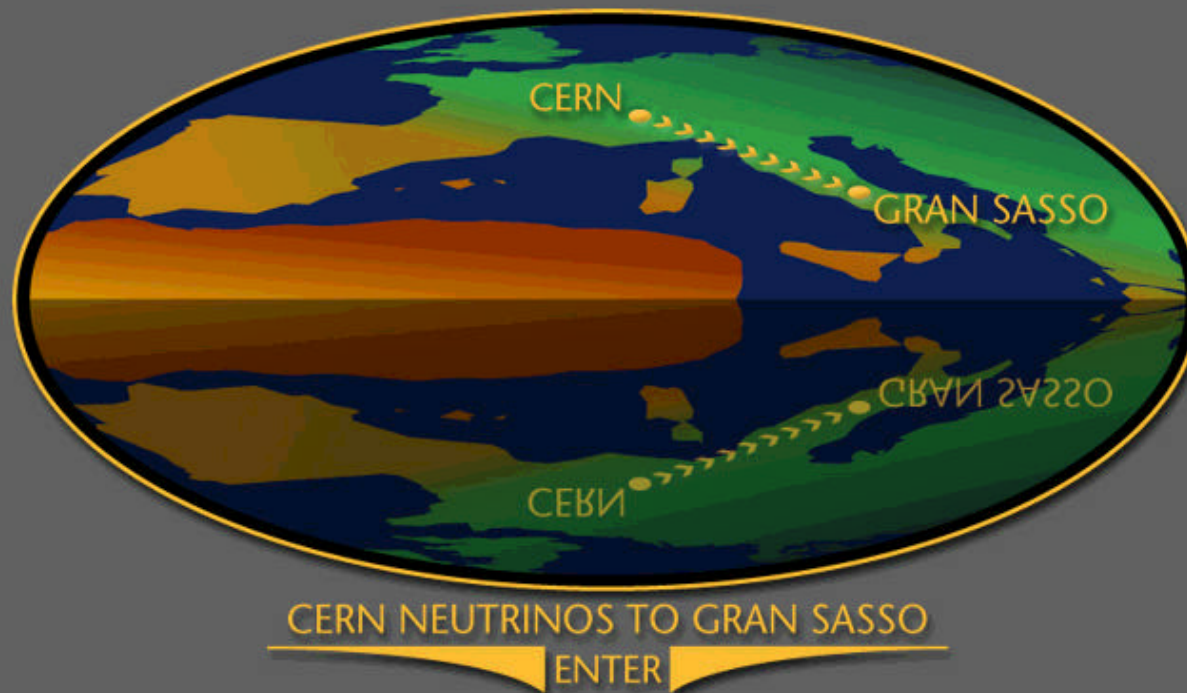


...before the summary...
for more information:



● CNGS general description

+ web pages



<http://proj-cngs.web.cern.ch/proj-cngs>



SUMMARY ... (the movie)