MAGNETIC HORNS for CERN neutrino beams from the sixties to 2003

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1. INTRODUCTION

CNGS neutrino beam project is entirely in AB's hands (NPA, TCL, EF, ECP, EP, $\rightarrow AB$) Only exception is the EP horn construction project. This peculiarity needs now to be corrected. Progressive transfer of activity from EP to AB is starting and should end in 2006.

2. TECHNICAL BASICS

2.1 Magnetic horn is a focusing device.

Invented in 1961 by S. Van der Meer for collection of charged particles emerging from a target. Used mainly for focusing

- Pion / kaon beams for neutrino generation
- Antiproton beams for p p⁻ storage rings



HORN FOCUSING PROPERTIES

Target produces charged particles : positively charged pions and kaons are emerging at various energies and angles.



Usually 2 horns are needed to produce a parallel wide band beam where a much larger number of particles emerging at various energies and angles are collected.



HORN/REFLECTOR - positive particle focusing of CNGS horn



F- Pietropaolo - INFN-Padova



CNGS focusing optics (positively charged particle trajectories)

Horn function in neutrino beam production summarized protons hit target (450 Gev in WANF & CNGS) π^+ / K⁺ produced in solid angle corresponding to acceptance magnetic horns to focus π^+ / K⁺ π^+ and K⁺ decay to μ^+ + ν_{Ω} in long evacuated pipe left-over hadrons shower in hadron absorber rock shield ranges out μ^+

 v_{O} beam travels through earth to experiment



Inner conductor profile

(current ranging from 100kA to 400kA)

Functional requirements for physics

- transparency of current sheet \rightarrow thin Al alloy sheet
 - → minimum supporting material within magnetic volume

- high magnetic field

 \rightarrow high current

Efficiency of horn system

Data from F. Pietropaolo – INFN / April 2003

CNGS reference beam

Nu_mu CC interactions (<10)0C	eV) at LNGS	5 (%)
Horn_Refl. zero-current	:	10%	
Reflector zero-current	•	~ 50%	
Standard configuration	:	100%	
Horn-Refl. half thickness	•	111%	
Horn-Refl. zero thickness	:	123%	
Perfect focusing	•	~ 200%	

2.2 How is a horn powered?

Horn is a short circuit i.e. small inductance 0.1 to 3 μH and small resistance 100 to 1000 $\mu \Omega$

Current is produced by discharge of a capacitor bank into horn in resonant mode (resonant L C R circuit)





High currents \rightarrow pulsed mode (reduced heating and power consumption) Pulse length can vary from the 10µs range up to the 10ms range. T/2 $\cong \pi$ (LC)^{1/2}







Typical horn capacitor discharge circuits



voltage reducing pulse transformer

Use of voltage reducing pulse transformer near horn is envisaged when distance between capacitor bank and horn has to be long. (WANF and CNGS)

V2 = V1/m

Simplifies also the thyristor switch (since I1 = I2/m)

Disadvantage :

- pulse length is longer by factor m

T/2 \cong m π (LC)^{1/2}

- addition of parasitic impedance

2.3 Operating voltage ?

Exposure to very high radiation doses

 \rightarrow insulation with mineral material and/or air

 \rightarrow preferably low voltage on horn (and busbars)

V (1) I (L/C)1/2

If heating is not critical, pulse can be long. Low voltage on horn is possible (i.e. < 1000V) :

-> high C

or -> voltage reducing transformer

Ex. LV horns : -low voltage capacitors without transformer ex. : NUMI at Fermi lab

> -high voltage capacitors with transformer ex. : WANF, CNGS, KEK

If heating is critical, short pulse lengths are compulsary. Pulse transformer cannot be used. HV on horn cannot be avoided.

- MiniBoone at Fermilab (4 kV range)

- Neutrino Factory prototype (4 kV range)
- PS neutrino beam with heavy liquid bubble chamber and Gargamelle in the 60's (12 kV range)

2.4 Flat top requirements



Particles cross magnetic volume when field is maximum. Duration of ejection fixes flat top requirement and pulse length needed.



2.6 Simulations

Electricity

Circuit is simple.

<u>But</u> main difficulty is correct evaluation of all inductive and resistive components. Few are negligible since horn is somehow a short-circuit. Skin effect has to be taken in account for short pulses. (For Al alloy 6082, δ = 1.6 mm for 2500Hz i.e. 200 µs pulse)



Requirement for CNGS horn : 95% probability to survive 2 10⁷ pulses

- Magnetic forces are cyclic $p = (\mu_0 I^2/8 \pi^2 R^2) (1-\tau/6)$ $\tau = t/R$ with t=wall thikness
- Thermal effects due to current are also cyclic with addition of beam energy deposition
- → Repetition rate must not hit a natural frequency of horn system
- Calculations are made using ANSYS
 - \rightarrow static FE analysis to estimate thermal stresses
 - \rightarrow modal analysis to estimate natural frequencies
 - \rightarrow ANSYS FE dynamic stress calculations
 - → fatigue analysis (evaluation of maximum equivalent completely reversed stress σ_{ecrs})
- → measurements on BA7 test circuit of natural frequencies and displacements due to vibrations

LAL-Orsay- S. Wallon CERN-EST-ME - Marek Kozien et al.

CNGS HORN INNER CONDUCTOR - 2 first modes determinant



LAL-Orsay- S. Wallon

CNGS HORN



J.M. Maugain EP-TA3

2.7 Summary of engineering constraints

Radiation	target - in front of - or into horn
	 → heating by particle absorption → beam effects on material strength (dislocations in metal)
	→ radiation resistance of materials → insulation problems
	\rightarrow access problems -remanent dose rates

Corrosive environment

Minimum thickness

light Al alloy inner conductor "10 kg robust Al alloy outer conductor" 400 kg

 Cyclic heat load
 by Joule effect

 by particle absorption
 > cyclic heat effect dilatation

 →cooling
 sprayed water along inner conductor

Cyclic magnetic forces

→ cyclic mechanical strain→ induced vibrations

 \rightarrow all effects produce mechanical fatigue

Pulsed electric system

<u>High current discharges</u>

 \rightarrow special capacitors

 \rightarrow thyristor assemblies

24 April, 2003.

Goal :

→ improve reliability and life time of horn
 → inner conductor as thin as possible
 ↓
 pushes mechanics close to the limit
 Fatigue is thus the limiting factor

Enormous progress on electrical part since thyristors replaced ignitrons

2.8 Handling problems linked to radiation

On site repairs are impossible due to high remanent dose rates. (10 mS intervention allowed/month/person,provided dose rate < 20 mS/h)



A Fast Coupling System is under development at LAL/Orsay which may relieve of the necessity to use a robot for CNGS horn replacement. Automatic procedures without human intervention should prevail.

3. 40 YEARS of HORN PROJECTS at CERN

3.1 PS neutrino beams from 61 -> 75

2s cycle with fast extraction from the PS of "2 μs proton burstat 25 GeV(up to 10E12 protons on target)



One of the first horns (63-64 PS neutrino experiment)









South East Area of PS










Logbook of 9 years neutrino runs (66-75)

Difficult runs (insulation problems, cable fire in 69, bad firing of ignitrons, bad electrical contacts,..) but big reward

Discovery at Gargamelle in 1973

The first observation of weak neutral current interactions was made in 1973 at the Gargamelle bubble chamber that was recording neutrino interactions at the CERN neutrino beam. (Gargamelle et les courants neutres – André Rousset Témoignage sur une découverte scientifique Collection Sciences de la Matière, 1996)

This was the first (indirect) evidence of the existence of the W and Z particles which were explicitly discovered 10 years later also at CERN by the UA1 and UA2 experiments (Carlo Rubbia and Simon van der Meer obtained Nobel prizes for this discovery).

End of runs in South West Area "1975

3.2 SPS neutrino beams from 77 to 98



fast/slow extraction from the SPS : two 6 ms long spills at 450 GeV (1 - 1.5)X10E13 protons on the target



Schematic Neutrino Beam Line





West Area Neutrino Facility (WANF)

Top view of neutrino cave





First horn in 76

Reflector positioning base







2 reflectors in serie with remote plug-in systems



CHORUS-NOMAD horn (110 kA) installed in 1993



CHORUS-NOMAD reflector (120 kA) installed in 1993



New TESLA transformer installed in 93 -transformer ratio 32-







Horn and reflector capacitor banks in BA7

Logbook of 21 years neutrino runs at WANF

High operational reliability

Current on horn kept low "110 kA

- to ensure run stability
- to avoid accidental horn replacement in very high radiation area

Several horn/reflector configurations have been operating for the succeeding group of experiments. No accidental horn exchange had to be faced after 77.

Complete refurbishment of target chamber took place in 92-93 for installation of CHORUS-NOMAD beam.

Impressive number of neutrino experiments based on WB beam, NB beam in 1984, "beam dump" beams, also ■ oscillation experiment at PS in 81-82.

3.3 CNGS to operate in 2006

Goal of the CNGS project (F- Pietropaolo - INFN-Padova)



"Long Base-Line" $v_{\mu} \rightarrow v_{\tau}$ oscillation experiments

- build an intense high energy ν_μ beam at <u>CERN-SPS</u>
- optimized for v_t appearance search at <u>Gran Sasso</u> laboratory (730 km from CERN)

CNGS Layout - F- Pietropaolo - INFN-Padova





Main design criteria of CNGS horns project

safety

savings maximum re-use of WANF equipment reliability



Horn electrical circuit layout for CNGS













BA7 horn test area (CNGS prototype testing in 2000-2001) horn with modified wanf inner conductor

















Recuperated WANF TESLA transformers -transformer ratio 16 (2&2 secondary couplings not mounted)

24 April, 2003.

Operation parameter table updated April 2003

Updated 15 April 2003	Unit	HORN SYSTEM	REFL. SYSTEM			
Duty cycle		2 pulses 50 ms apart all 6 s				
Peak current in horn	kA	150	180			
Transformer ratio		16	32			
Primary current peak	kA	9.375	6.646			
Total capacitance for two switching sections	μF	45.4 x 90 x 2 = 8172	45.4 x 90 x 2 = 8172			
Pulse duration	ms	7.5	10			
Charging voltage	V	7700	6300			
Total stored energy	kJ	$2 \times 119 = 238$	$2 \times 80 = 160$			
Max. voltage on element	V	280	150			
Mean power dissipated in element	Lw	16	10.5			
by current only (2 pulses)	K VV	10	10.5			
Mean power dissipated in element	kw	10	6			
(inner + outer conductor) for $7.2 \times 10^{**13}$ pot per 6s cycle						
Total power dissipated in element (2 pulses)	kw	26	16.5			
Waterflow for $\delta\theta$ out- $\delta\theta$ in = 5°C	l/min	75	48			

Status of CNGS horn project in April 2003

	20	00	2001	200	2	2003 2004		2005		2006		
Production of horns at Orsay												
New capacitors												
Cable order												
Reconditioning of electric circuits												
Testing of LAL horns												
Installation in BHA4/ECA4												
Installation in target chamber												
Commissioning and tests												
3.5 Neutrino Factory studies for future

Prototype developed at CERN in 2001

Maximum current	:	300 kA
Pulse repetition rate:		50 Hz
Waist diameter	:	80 mm
Length	•	1030 mm

Life time

Pulse length r.m.s current Voltage on horn

Skin depth Joule losses 6 weeks or 2 x 10^8 pulses for the prototype (six months or 8 x 10^8 pulses for a final horn) : $\leq 100 \ \mu s$: $15 \ kA \ (CMS - 20 \ kA)$: $\sim 4000 \ V$

: 1.25 mm

Main technical problems : Cooling, vibrations, irradiation, fatigue



Studies of NUFACT horn are taken over by LAL/Orsay in the frame of European Neutrino Group created in 2003



CERN prototype under test

4 CNGS HORN PROJECT TODAY

CERN groups AB/PO with the support of AB/CO AB/ATB EP/TA3 collaborate in the frame of the newly created

Joint Horns & their Power supplies Construction project for CNGS (JHPC-CNGS)

EP Division is responsible to deliver the horn systems in 2006 as turnkey systems to AB division.

• AB division will take over full responsability for the future operation and maintenance.

Joint interdivisional efforts and AB standardised approaches will help to ensure a smooth transition to the operation phase.

Production of horns including water cooling systems, striplines and supports has been outsourced to IN2P3 - LAL/Orsay as inkind contribution according to :

MoU of 4 August 2000 between CERN and IN2P3 - LAL/Orsay

LAL/Orsay will :



Produce horn/reflector and striplines including water circuits and supports as inkind contribution.

(Outsourcing organised by EP/TA3)



Take part in electrical testing, final installation and running-in.

5. SUMMARY

Magnetic horns are a fundamental instrument in the generation of accelerator neutrino beams.



6. CONCLUSION

What has changed in 40 years?

- electronbeam welding, less supporting material
- power thyristors
- computer controls
- elaborate calculations with computer programs
- distance of detectors

How does future look ?

AB division goes into a nice heritage in 2006. Let's wish that neutrino hunting goes on successfully with the expected \blacksquare_{τ} appearance (τ lepton detection).

On the longer future, development of fast cycling horns for neutrino factories opens a new field of developments. (neutrino oscillations proposals with a Superbeam and Beta-beam ?) **Problems with neutrinos ??...** (see recent novel from F. Vanucci)

Many thanks to all from the sixties to now